

天关对恒星 X 射线耀发大样本观测 的初步结果

报告人：毛选

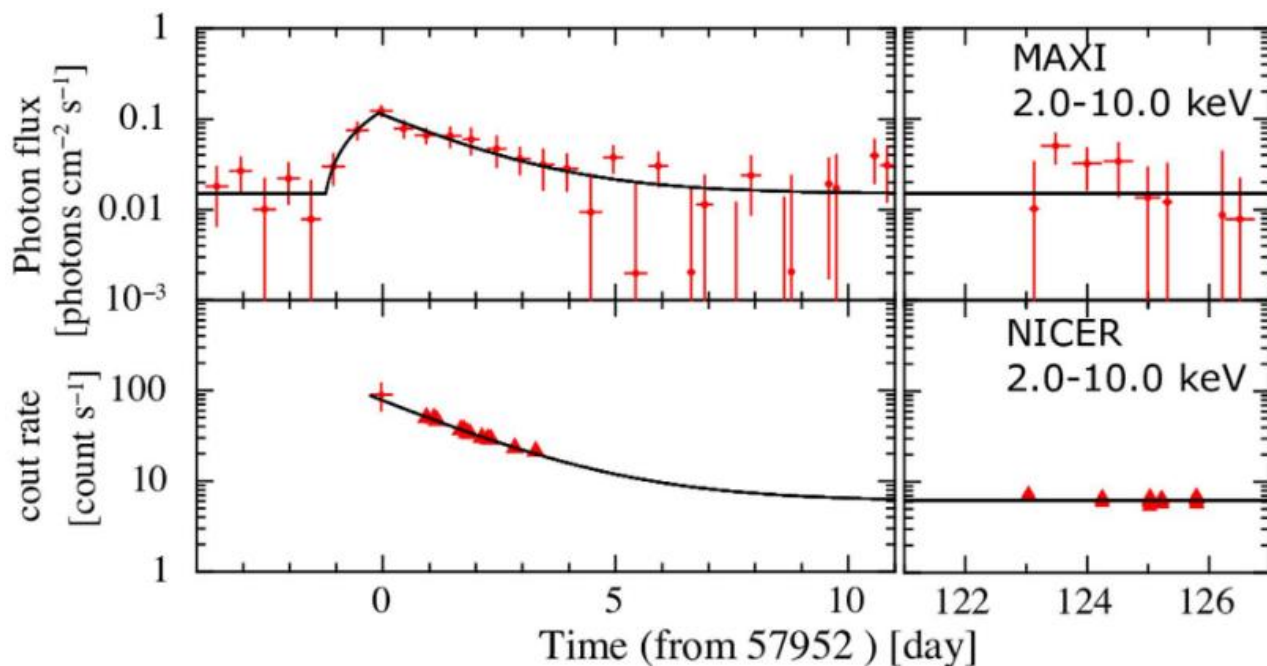
单位：国家天文台

2025年5月27日

EP之前的恒星X射线耀发探测：个源研究

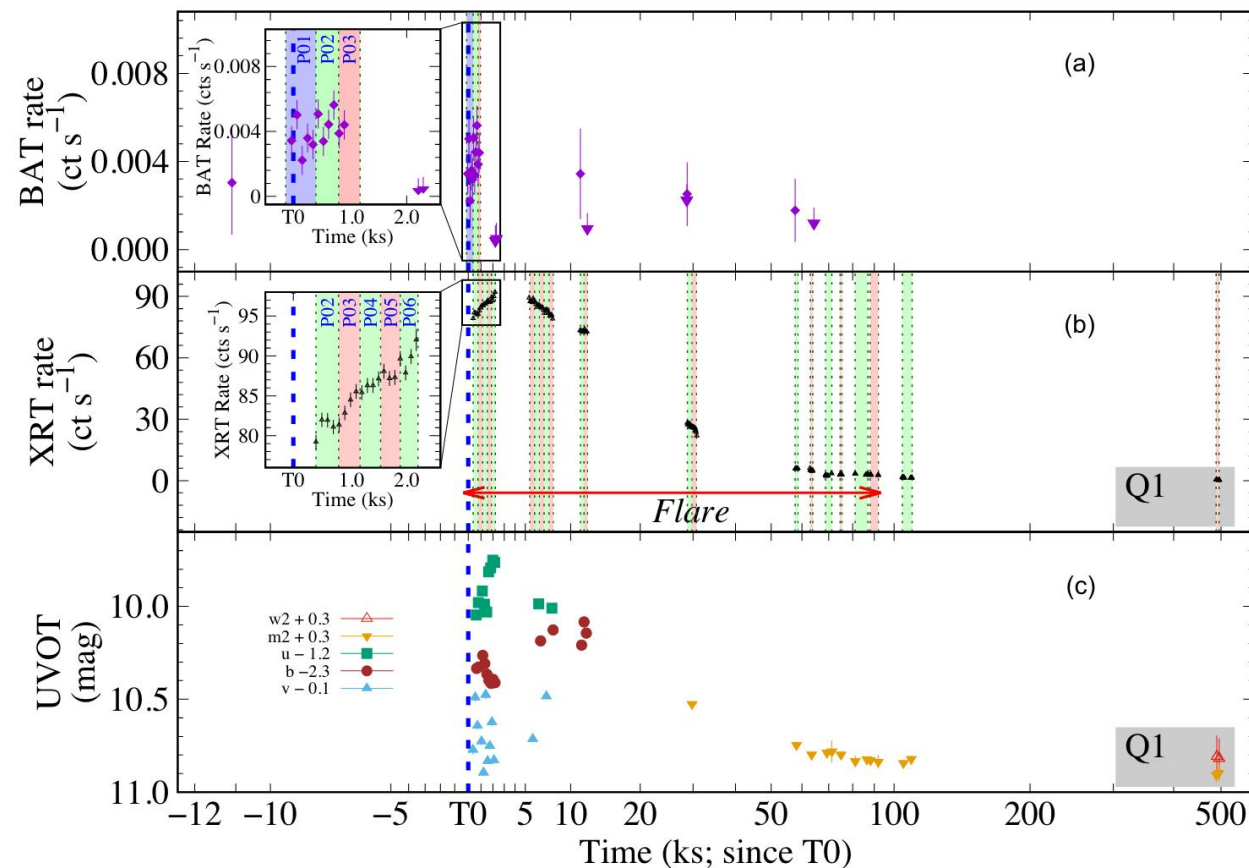


- MAXI/Gas Slit Camera(GSC)



Sasaki+ 2021

- Swift/Burst Alert Telescope(BAT)



Karmakar+ 2022

EP之前的恒星X射线耀发探测：样本研究

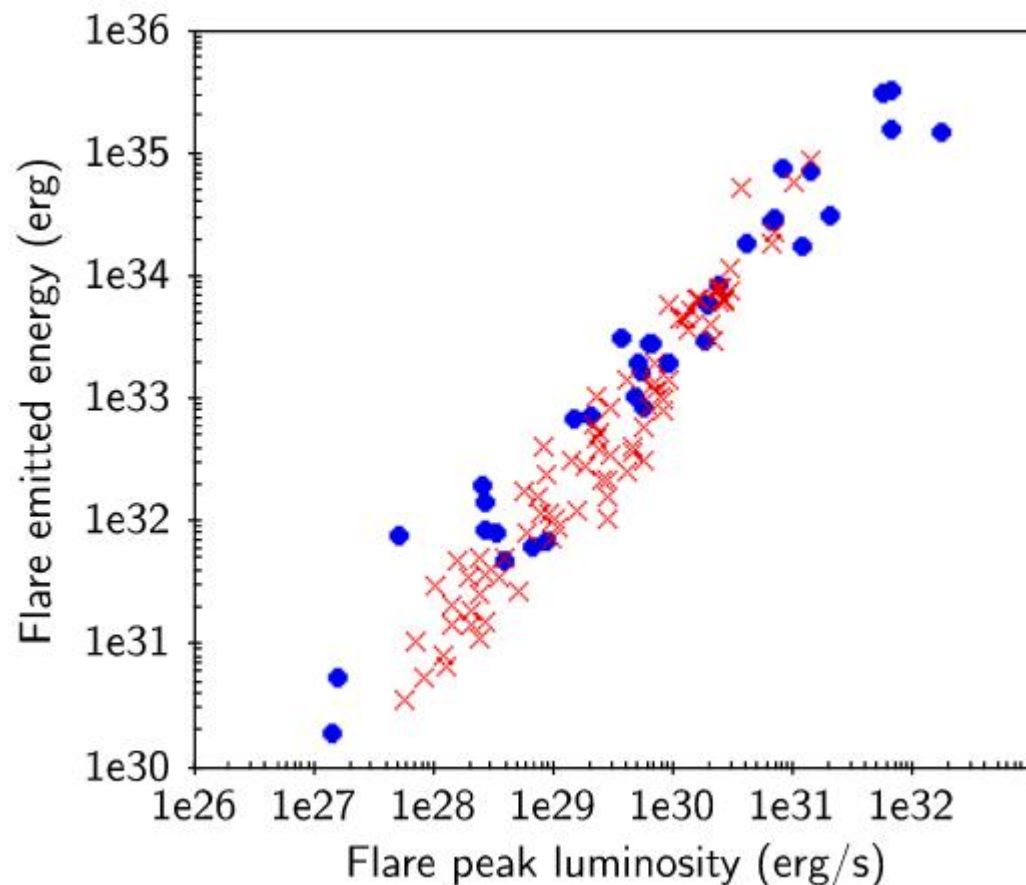


1. 巡天中的偶发探测

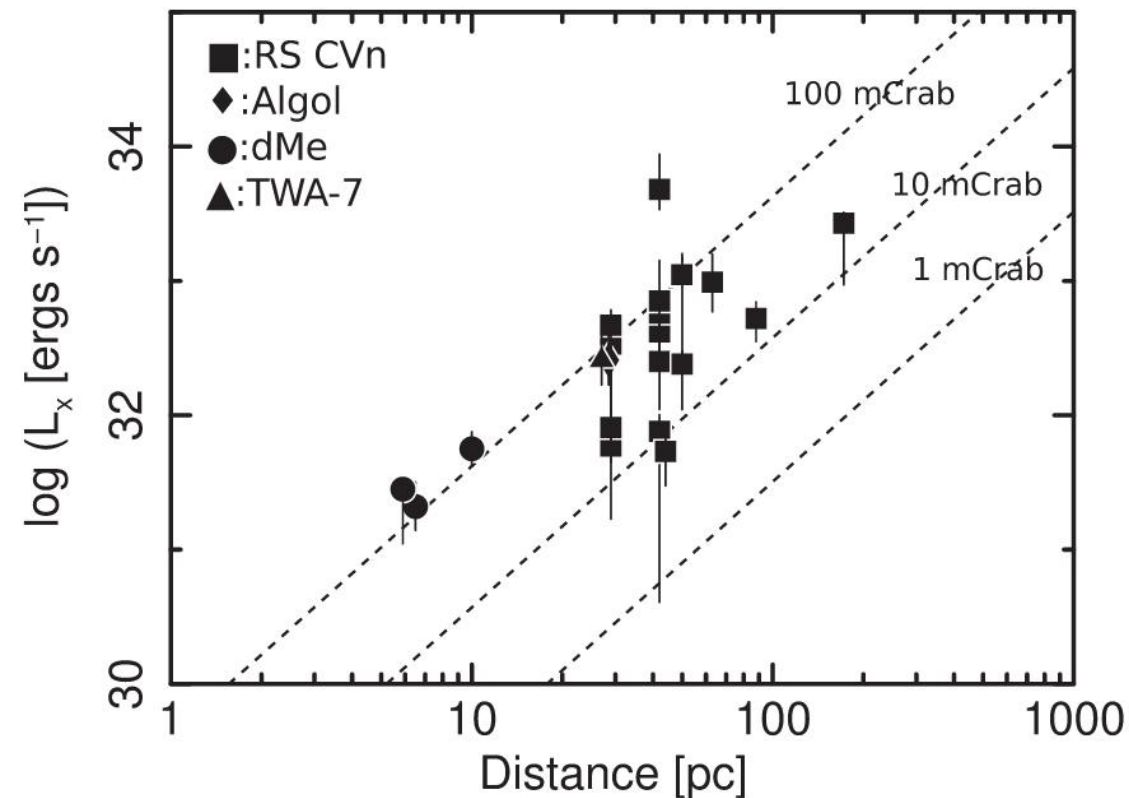
Variability type	Sample	Number of stars	Number of light-curves	Number of events	
				Total	Completely observed
Flare	CVS	63 (30)	76 (32)	133 (39)	116 (34)
	CLVS	8 (4)	8 (4)	11 (4)	11 (4)
	CVS+CLVS	70 (34)	84 (36)	144 (43)	127 (38)
Trend	CVS	10 (1)	11 (1)	11 (1)	
Gradual	CVS	3 (1)	3 (1)	3 (1)	
	CLVS	1 (0)	2 (0)	2 (0)	
Indeterminate	CVS	6 (5)	6 (5)	6 (5)	
	CLVS	2 (0)	2 (0)	2 (0)	
Poor bgd sub? ^a	CVS	2 (0)	2 (0)	2 (0)	
All	CVS	76 (36)	96 (39)		
	CLVS	11 (4)	12 (4)		
	CVS+CLVS	84 (40)	108 (43)		

Pye+ 2015

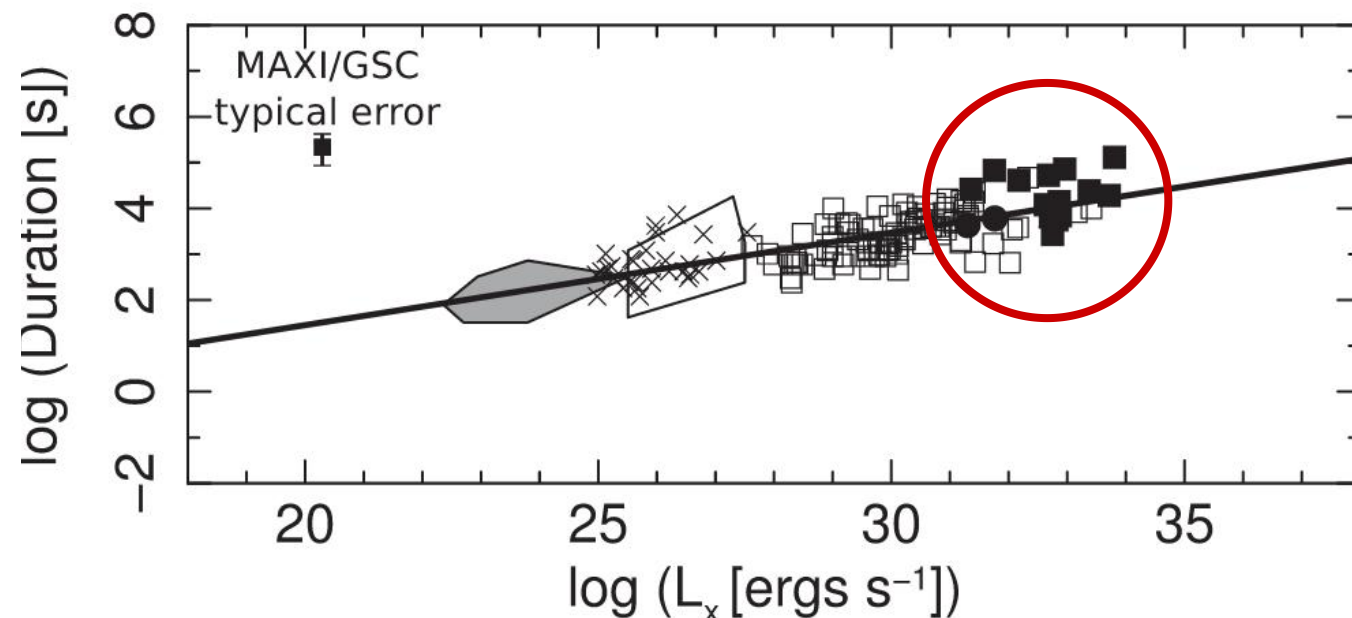
- ~70颗恒星
- ~130次耀发



2. 全天监测设备的探测



- 2年间探测到23次耀发，峰值 $10^{31} - 34 \text{ erg/s}$ (2-20keV)



Tsuboi+ 2017

EP之前的X射线恒星耀发探测：样本研究

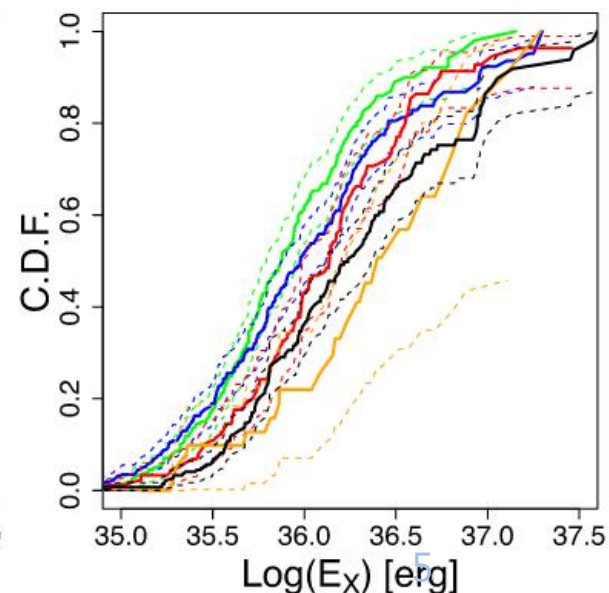
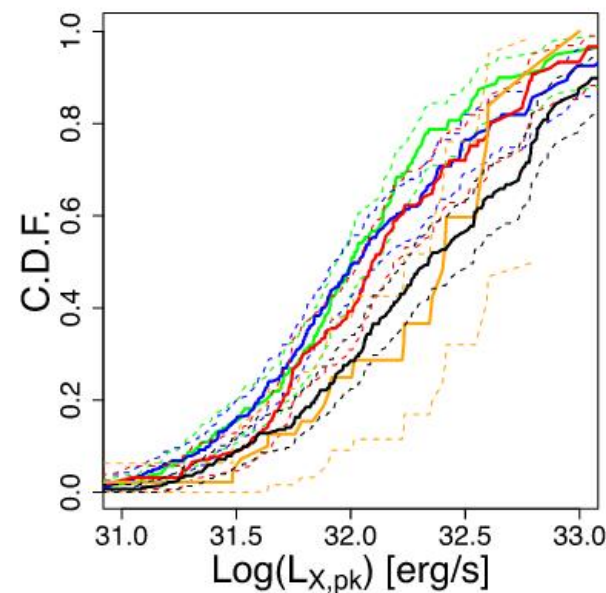
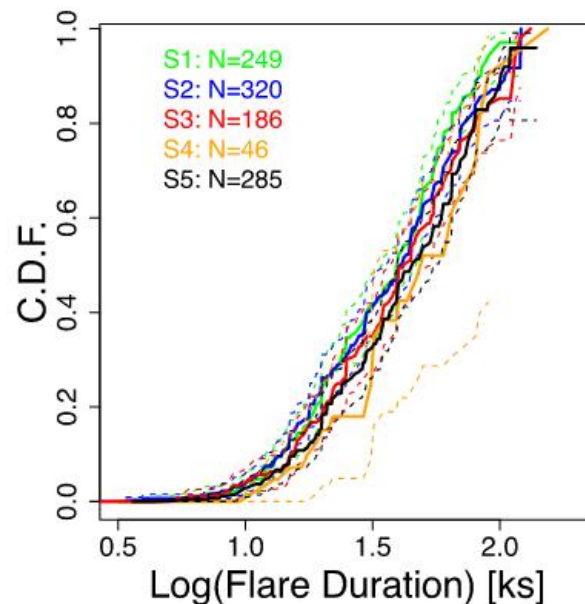


3. 恒星密集区的探测

Table 1
MYStIX and SFiNCs Regions

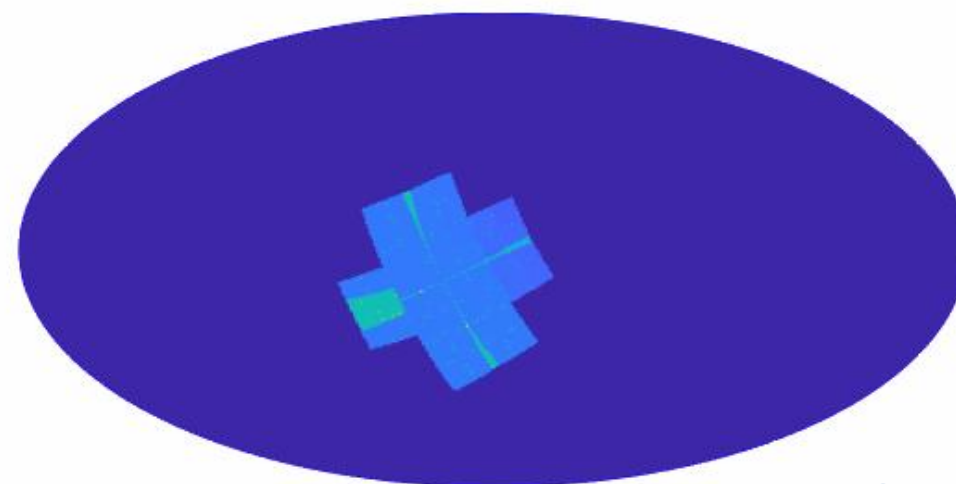
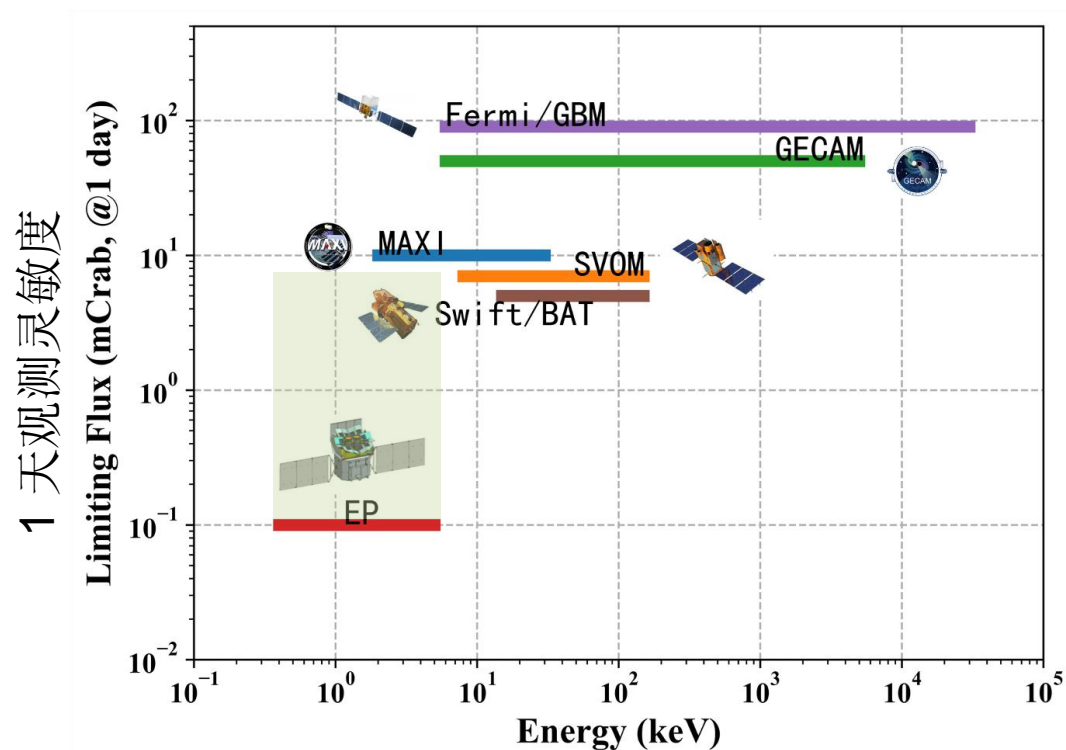
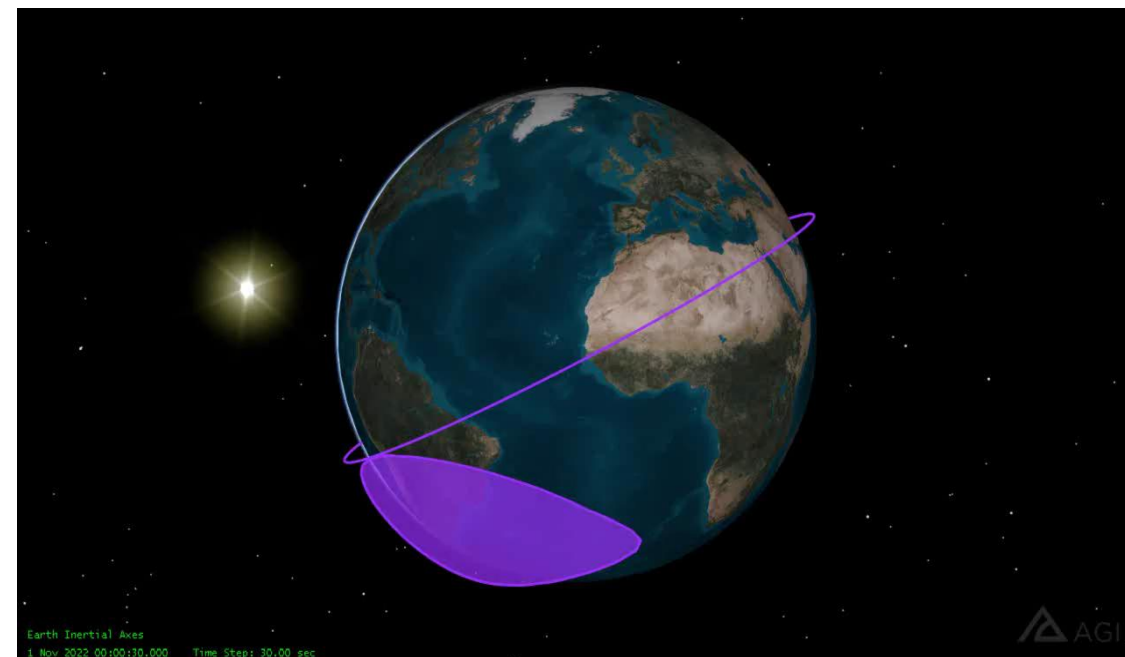
Region	R.A. (deg)	Decl. (deg)	Dis. (pc)	$N_{X,obs}$	N_{tot}	Region	R.A. (deg)	Decl. (deg)	Dis. (pc)	$N_{X,obs}$	N_{tot}
(1)	(2)	(3)	(4)	(5)	(6)	(1)	(2)	(3)	(4)	(5)	(6)
Be 59	0.6	67.4	1100	464	1703	NGC 2264	100.3	9.6	738	898	1837
BRC 2	1.0	68.5	1100	42	142	NGC 2362	109.7	-25.0	1332	467	512
Carina Neb.	161.2	-59.7	2620	6751	37899	NGC 3576	168.0	-61.2	2800	1131	12235
Cep A	344.1	62.0	868	194	534	NGC 6334	260.1	-35.9	1770	1385	12661
Cep B	343.9	62.6	868	1032	1773	NGC 6357	261.4	-34.3	1770	1952	12581
Cep C	346.5	62.5	868	95	247	NGC 7160	328.5	62.6	961	134	157
DR 21	309.8	42.3	1500	594	3907	OMC 2-3	83.9	-5.1	390	287	402
Eagle Neb.	274.7	-13.8	1740	2065	7762	ONC Flank N	83.8	-4.8	392	198	230
Flame Neb.	85.4	-1.9	414	422	596	ONC Flank S	83.8	-5.7	395	223	261
GGD 12-15	92.7	-6.2	830	141	300	Orion Neb.	83.8	-5.4	405	1414	1700
IC 348	56.1	32.1	324	307	284	RCW 120	258.1	-38.5	1680	262	1570
IC 5146	328.4	47.3	783	161	272	RCW 36	134.9	-43.8	930	337	1228
IRAS 20050	301.8	27.5	700	213	308	RCW 38	134.8	-47.5	1700	813	5750
LDN 1251B	339.7	75.2	300	39	39						
Lagoon Neb.	271.0	-24.4	1336	1828	5058						
Lkh α 101	67.5	35.3	564	197	311						
M17	275.1	-16.2	1680	2296	13412						
Mon R2	91.9	-6.4	948	410	882						
NGC 1333	52.3	31.3	296	116	111						
NGC 1893	80.7	33.4	3790	1110	5977						
NGC 2068	86.7	0.1	414	231	335						

- 主序前星的1086次耀发



EP观测模式

- * 轨道
 - ★ 高度592 km, 轨道周期 96min
- * 巡天模式
 - ★ 每轨3个指向, 每个指向 ~20min
 - ★ 3轨覆盖夜天区 (约5个小时)
 - ★ 半年覆盖全天



耀发恒星列表



Stellar Flare List

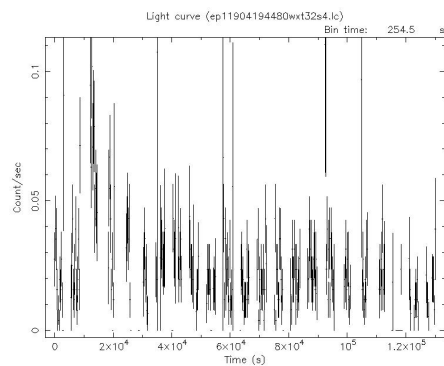
[Export CSV](#)

▲	Common Name ▲	RA ▲	Dec ▲	Pos Err ▲	Alert Time (UTC) ▲	Goto	Detail	Tags ▲	Category ▲	CL ▲	R_Flux ▲
1	BD+44 2051	166.369	43.527	0.015	2024-12-16 23:59:33	☉	ⓘ	known_source,stellar_flare	Star	High Proper Motion Star	3.12e-12
2	HD 028907	69.076	55.027	0.02	2024-03-01 10:11:10	☉	ⓘ	known_source,stellar_flare	Star	Spectroscopic Binary	6.42e-12
3	LP 705-28	6.462	-9.961	0.042	2024-07-29 18:02:32	☉	ⓘ	known_source,stellar_flare	Star	Eruptive Variable	
4	PM J17581-2222	269.526	-22.373	0.037	2024-09-17 07:40:43	☉	ⓘ	known_source,stellar_flare	Star	Rotating Variable	
5	GR Mus	194.405	-69.289	0	2025-05-16 05:53:03	☉	ⓘ	stellar_flare	XRB	LMXB	3.81e-12
6	HD 181321	290.374	-34.983	0.014	2025-05-11 00:20:11	☉	ⓘ	known_source,stellar_flare	Star	Spectroscopic Binary	4.16e-11
7	StKM 1-2018	338.344	-9.615	0.01	2024-09-12 04:29:07	☉	ⓘ	known_source,stellar_flare	Star	Double or Multiple Star	8.74e-12
8	BS Ind	320.249	-52.478	0.005	2025-05-23 00:14:34	☉	ⓘ	known_source,stellar_flare	Star	Spectroscopic Binary	9.67e-12
9	TYC 3904-65-1	274.031	54.13	0.04	2024-11-03 00:04:59	☉	ⓘ	known_source,stellar_flare	Star	Star	1.21e-11
10	V1362 Tau	80.543	24.536	0.018	2024-01-26 00:41:33	☉	ⓘ	stellar_flare,known_source	Star	RS CVn Variable	3.92e-12
11	HD 22468	54.197	0.588	0	2024-11-25 03:08:02	☉	ⓘ	known_source,stellar_flare	Star	RS CVn Variable	3.72e-12
12	EP250207a	356.902	27.027	0.037	2025-02-07 15:48:27	☉	ⓘ	known_source,stellar_flare	Star	Star	
13	PM J02146+3121	33.66	31.363	0.041	2024-11-26 03:04:02	☉	ⓘ	known_source,stellar_flare	Star	High Proper Motion Star	3.33e-12
14	V335 Peg	343.151	10.22	0.019	2024-12-04 09:28:47	☉	ⓘ	known_source,stellar_flare	Star	Spectroscopic Binary	1.34e-11
15	UPM J1217-3644	184.448	-36.747	0.039	2025-02-06 20:16:56	☉	ⓘ	known_source,stellar_flare	Star	High Proper Motion Star	4.47e-13

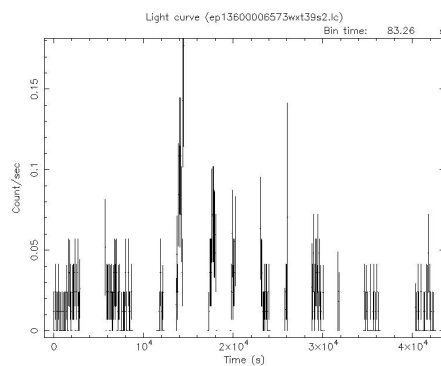
Showing 1-15 of 1134 rows

[First](#) [Prev](#) [1](#) [2](#) [3](#) [4](#) [5](#) [Next](#) [Last](#)

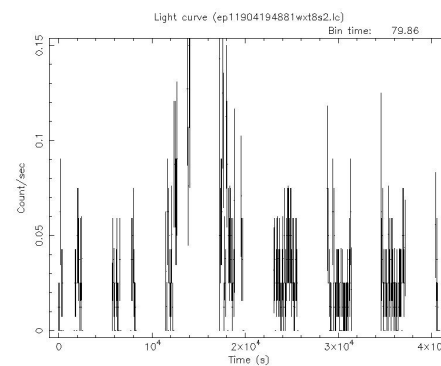
单次观测光变曲线中的恒星耀发



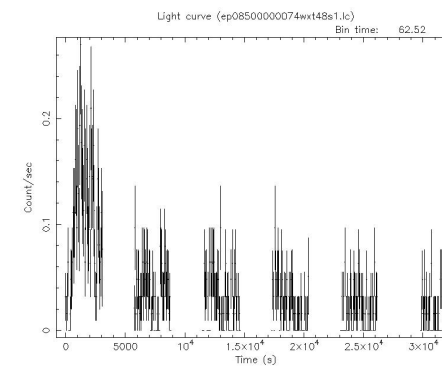
Start Time 20473 14:22:20:288 Stop Time 20475 2:29:54:485



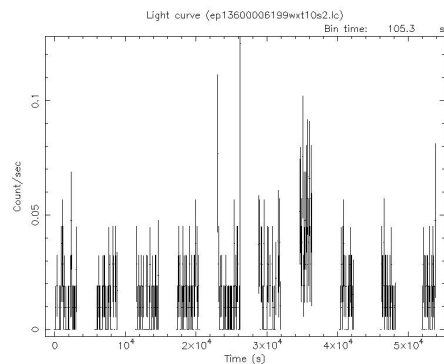
Start Time 20560 18:34:05:662 Stop Time 20561 6:23:09:510



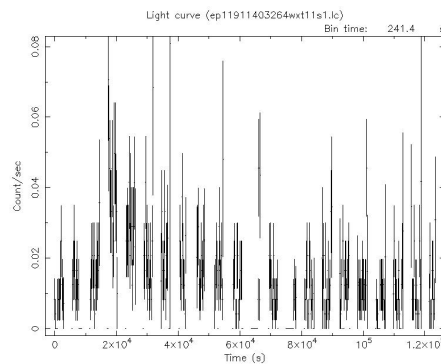
Start Time 20522 0:11:37:964 Stop Time 20522 11:31:46:020



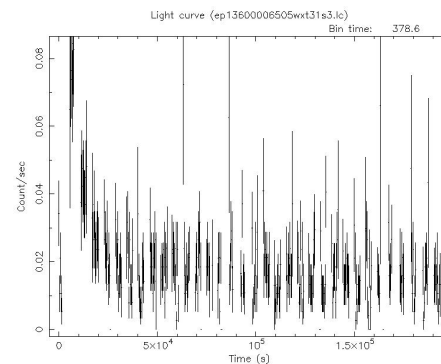
Start Time 20421 10:17:52:345 Stop Time 20421 19:10:21:286



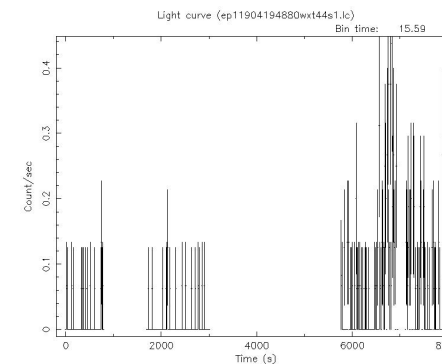
Start Time 20449 19:04:18:707 Stop Time 20450 10:01:30:569



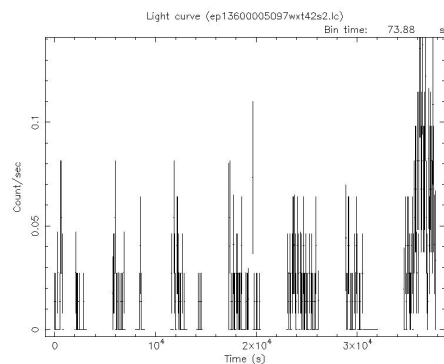
Start Time 20397 19:35:54:668 Stop Time 20399 5:51:34:121



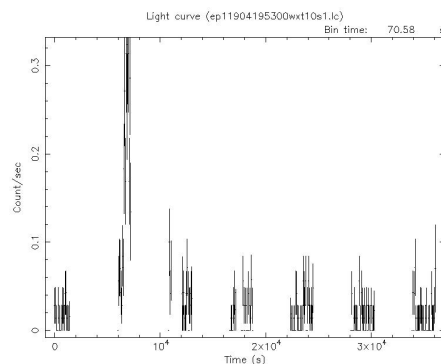
Start Time 20496 8:15:22:404 Stop Time 20498 14:00:07:898



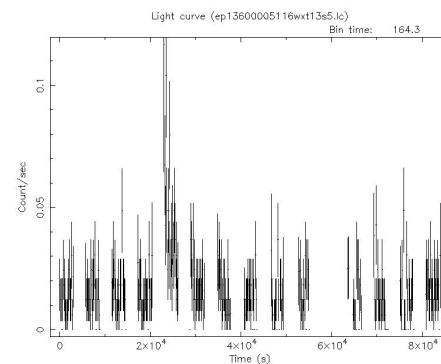
Start Time 20521 5:08:13:781 Stop Time 20521 7:21:02:543



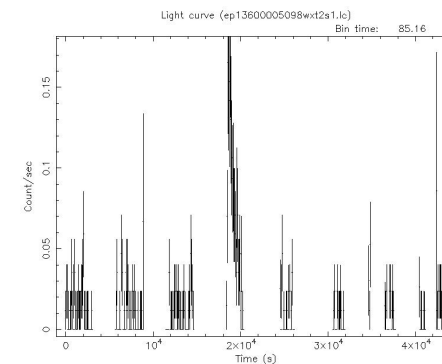
Start Time 20375 1:27:19:973 Stop Time 20375 11:56:31:422



Start Time 20580 19:52:57:634 Stop Time 20581 5:54:04:992



Start Time 20381 8:35:32:172 Stop Time 20382 7:54:37:612

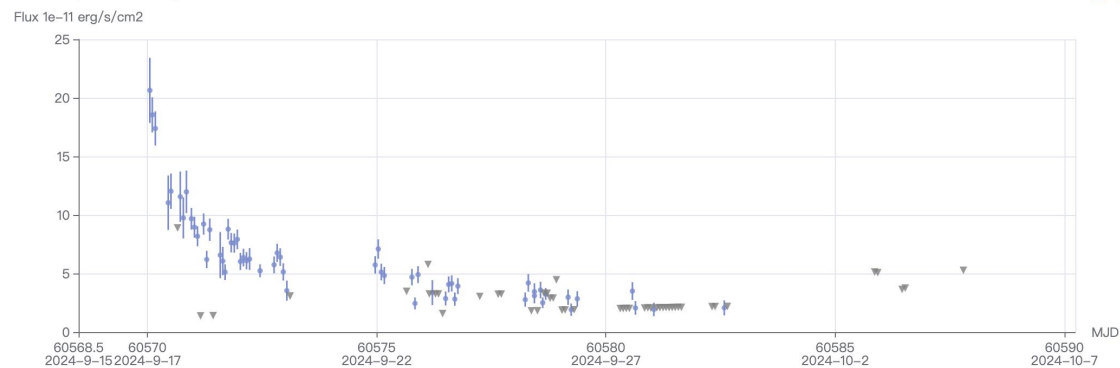


Start Time 20375 12:38:10:613 Stop Time 20376 0:43:26:550

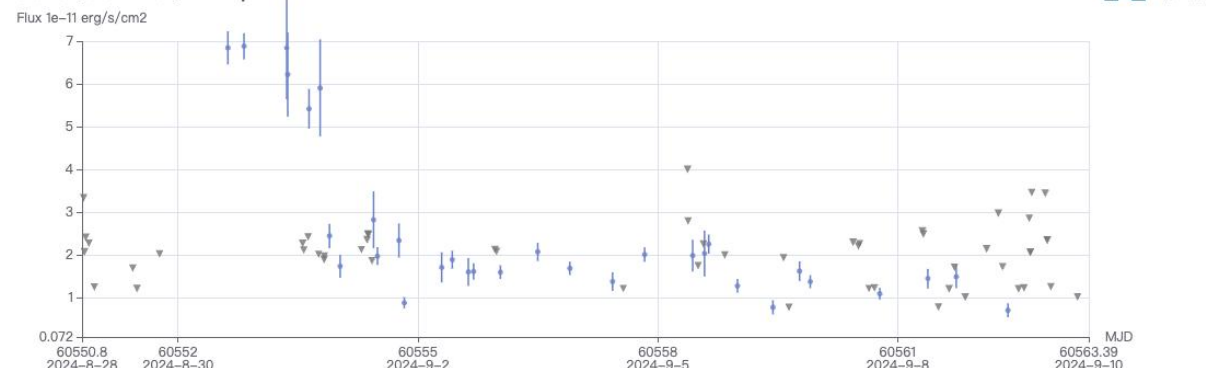
长期光变曲线中的恒星耀发



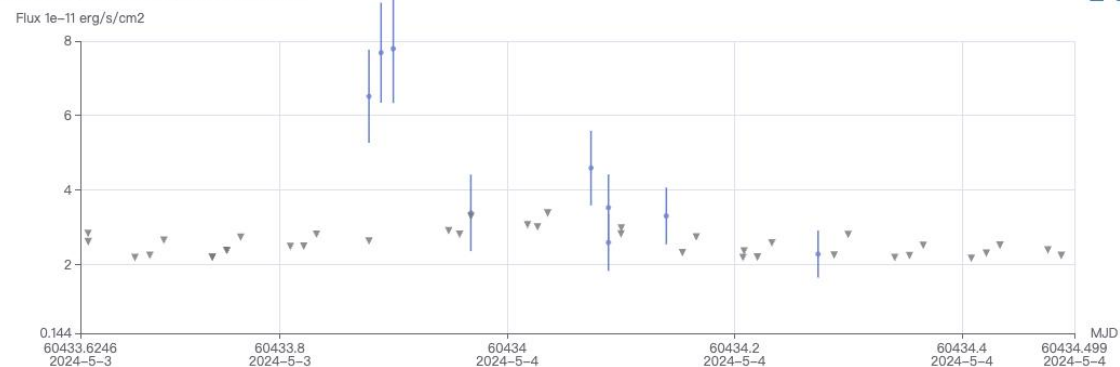
Light Curve (0.5–4 keV) 29 Dra



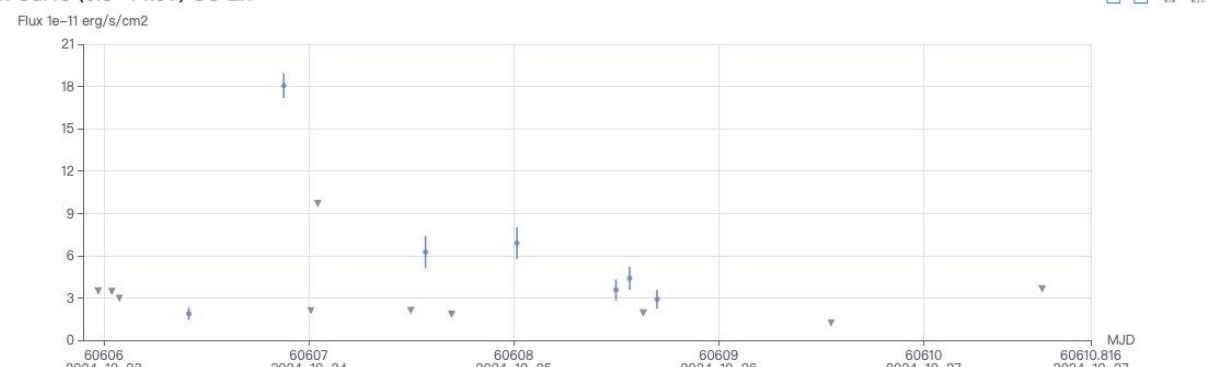
Light Curve (0.5–4 keV) FF Aqr



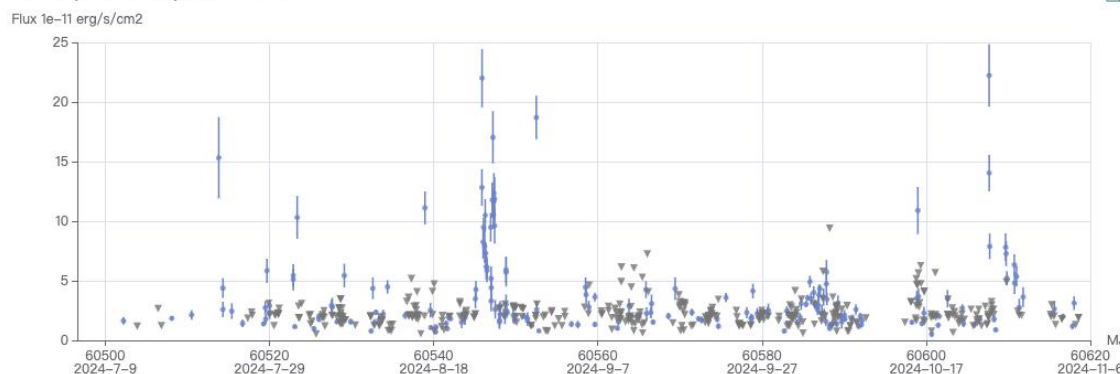
Light Curve (0.5–4 keV) BD+08 3048



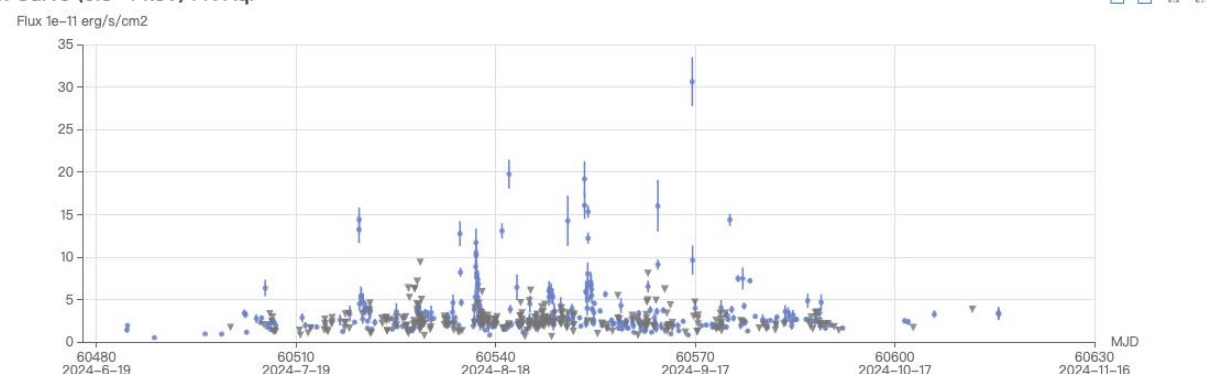
Light Curve (0.5–4 keV) CC Eri



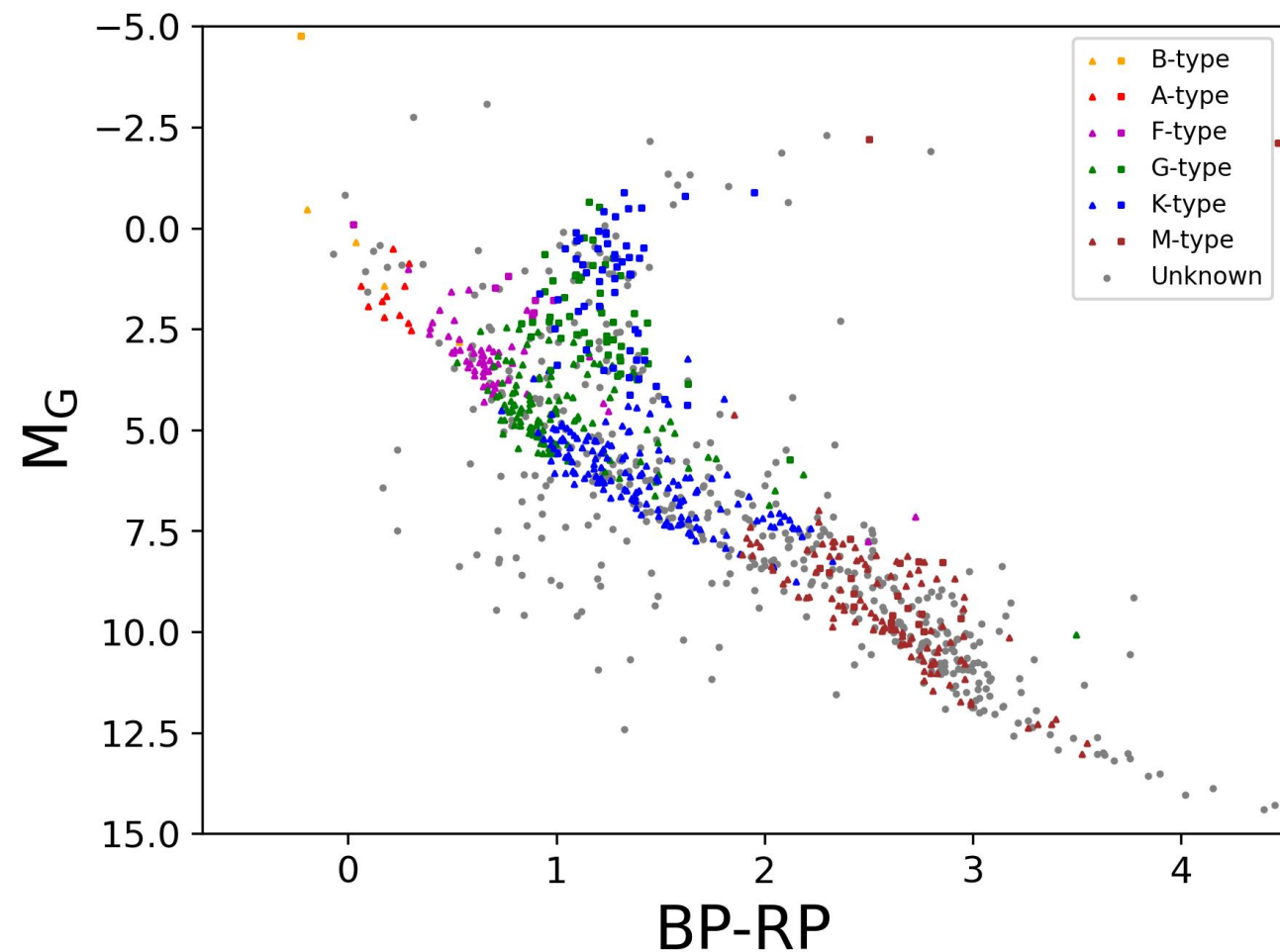
Light Curve (0.5–4 keV) BD+19 5116



Light Curve (0.5–4 keV) FK Aqr



H-R Diagram



总 计 : **1101 stars**
 有距离和测光数据 : 1036 stars
 有温度和logg信息 : 577 stars

- 平均每天探测到**2-3颗**恒星的耀发!

	O-type	B-type	A-type	F-type	G-type	K-type	M-type
T	>30000	10000-30000	7500-10000	6000-7500	5200-6000	3700-5200	<3700

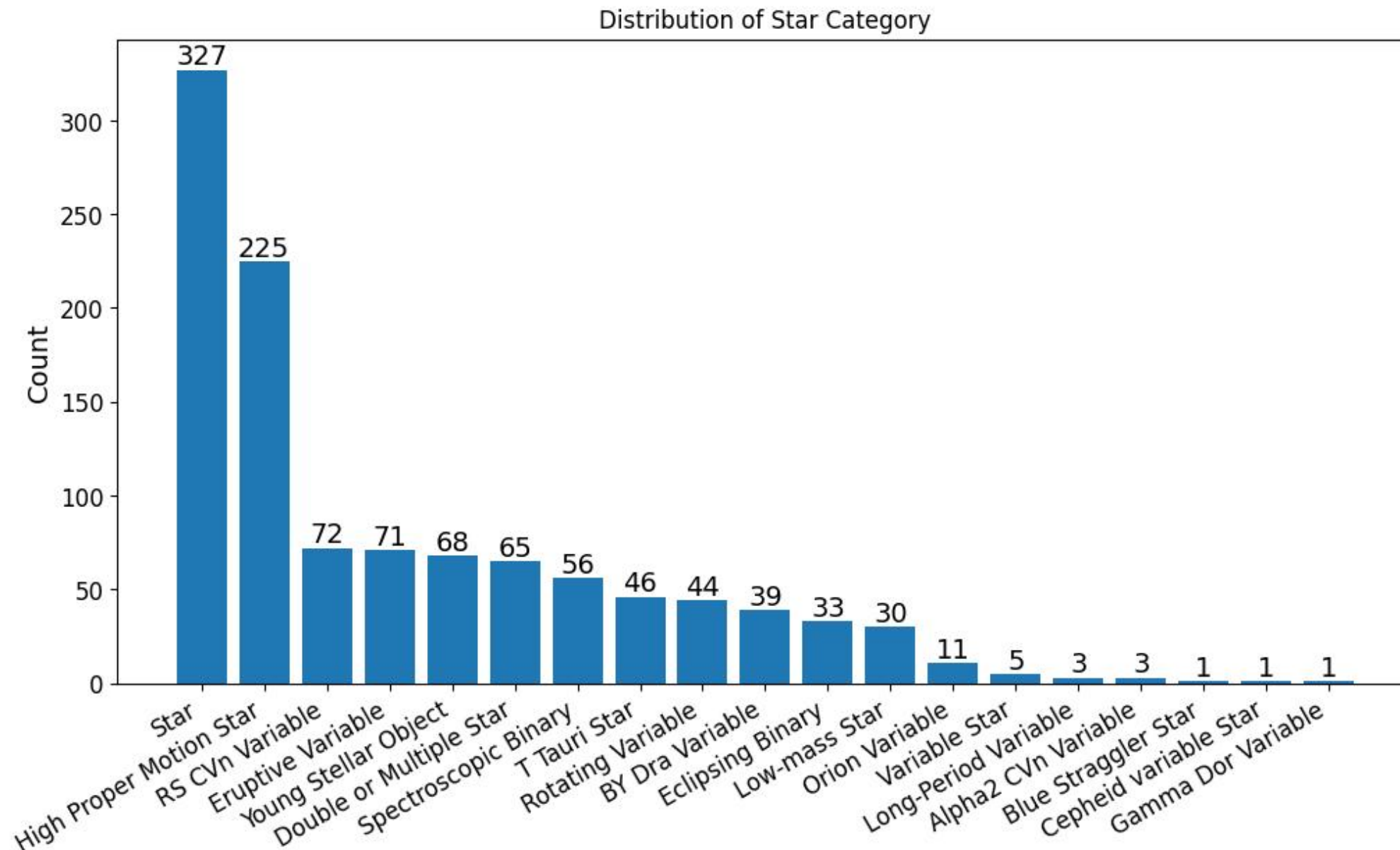
Giant:

$$\log g < \begin{cases} 3.5 & \text{if } T_{\text{eff}} \geq 6000, \\ 4.0 & \text{if } T_{\text{eff}} \leq 4250, \\ 5.2 - 2.8 \times 10^{-4} T_{\text{eff}} & \text{if } 4250 \leq T_{\text{eff}} \leq 6000. \end{cases}$$

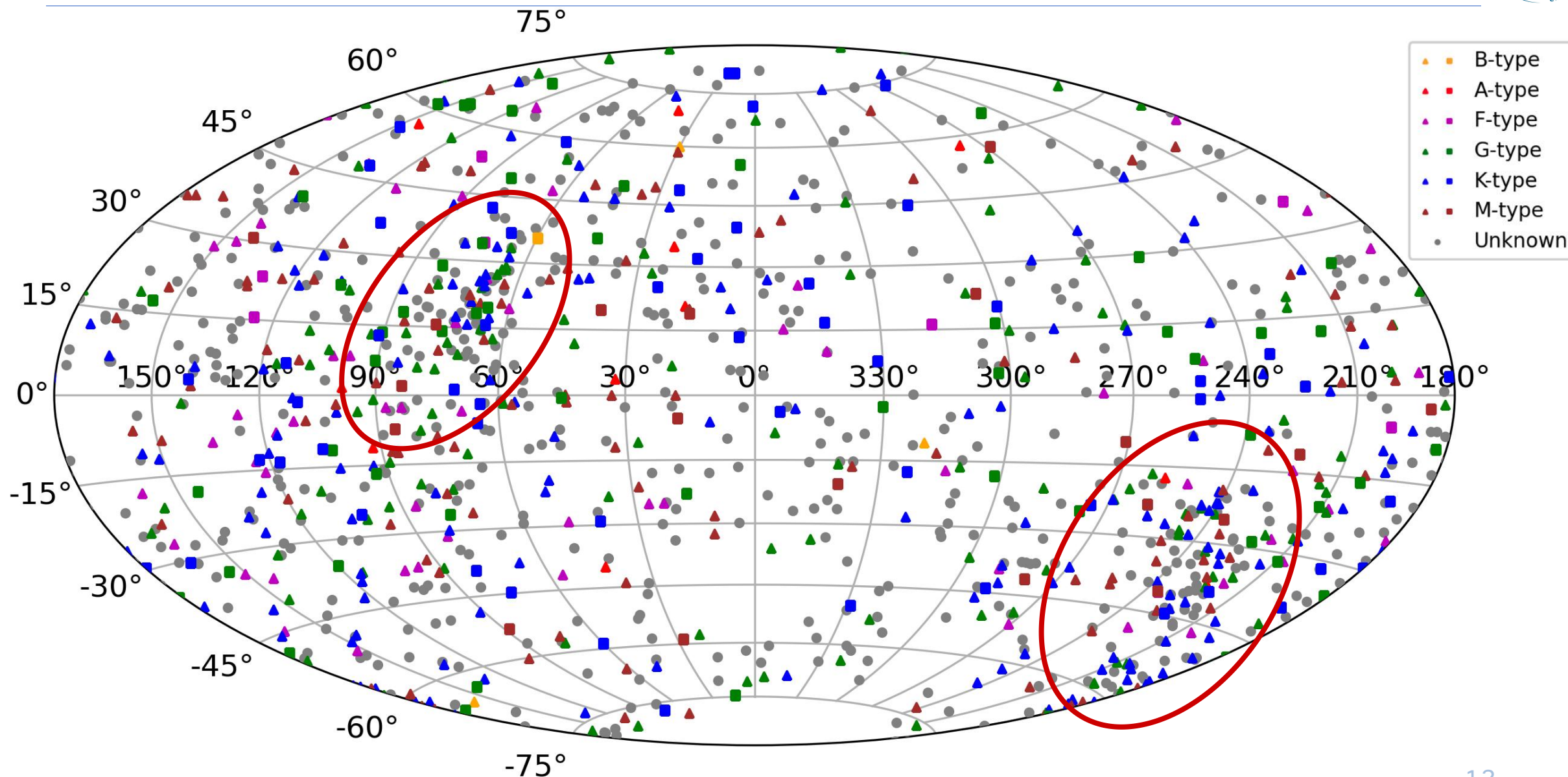
Giardi+ 2011

	O-type	B-type	A-type	F-type	G-type	K-type	M-type	SUM
Dwarf /YSO	0	4	11	58	118	148	106	445
Giant	0	1	0	7	51	54	19	132
SUM	0	5	11	65	169	202	125	577

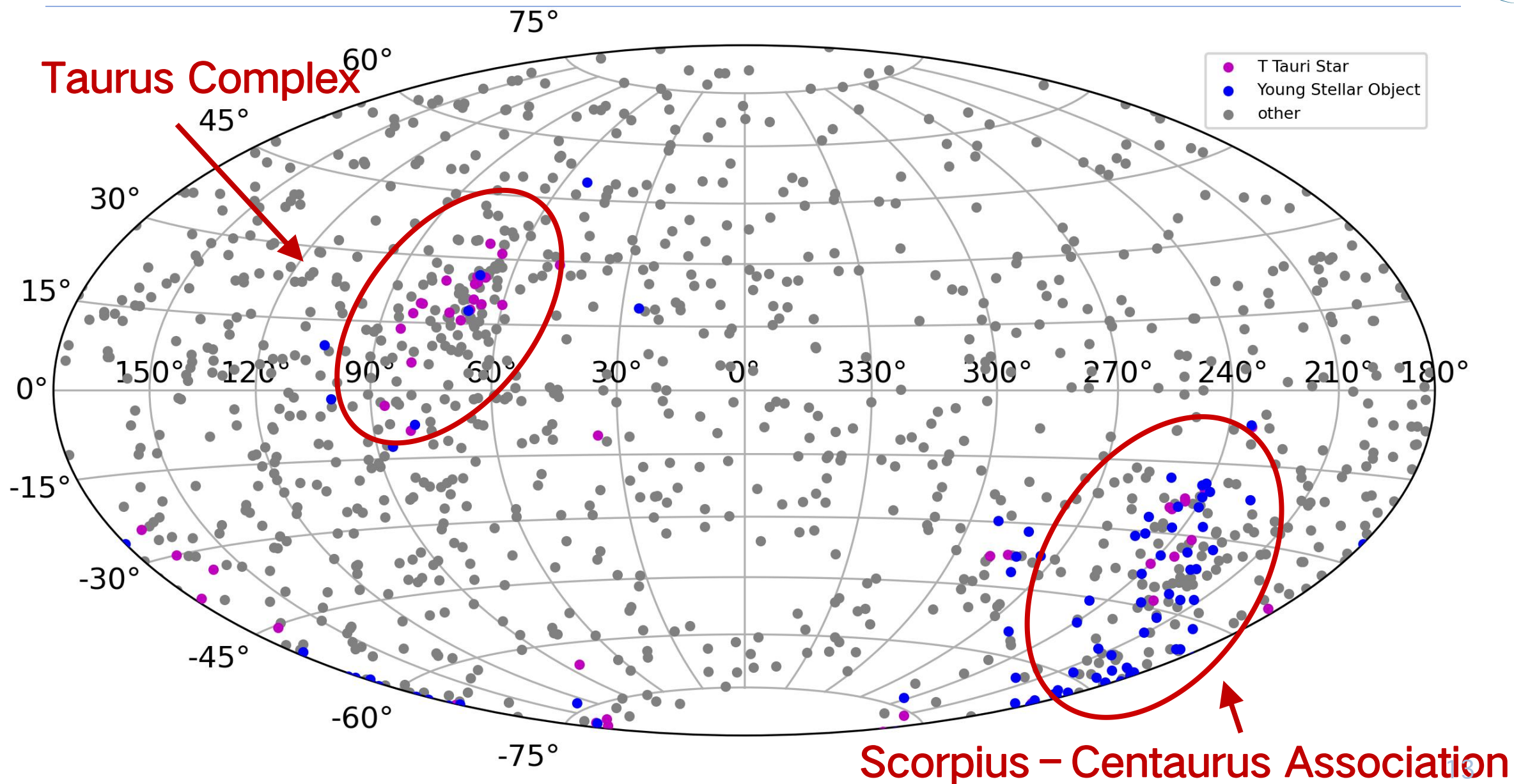
基于Simbad的耀发恒星分类



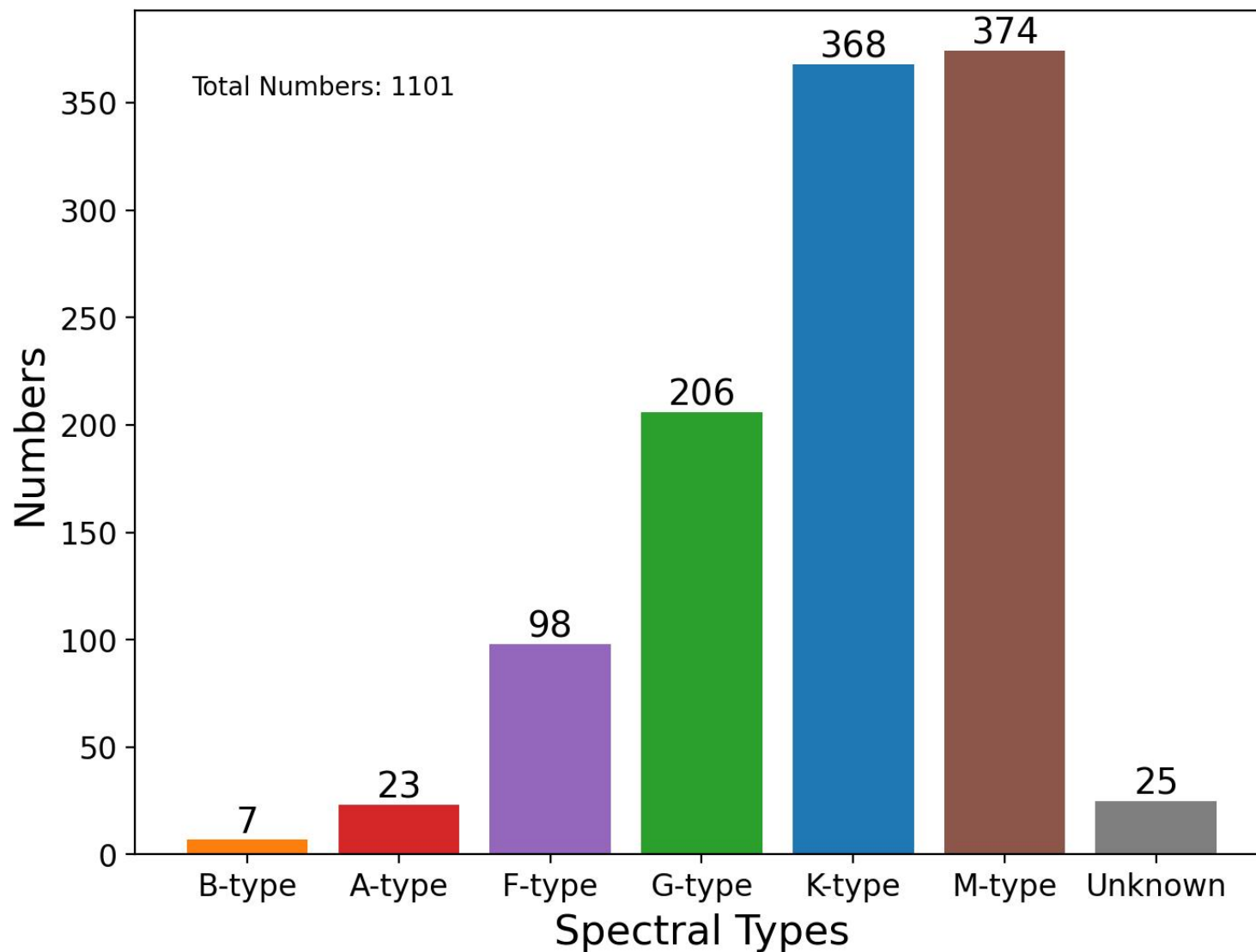
耀发恒星全天分布



耀发恒星全天分布

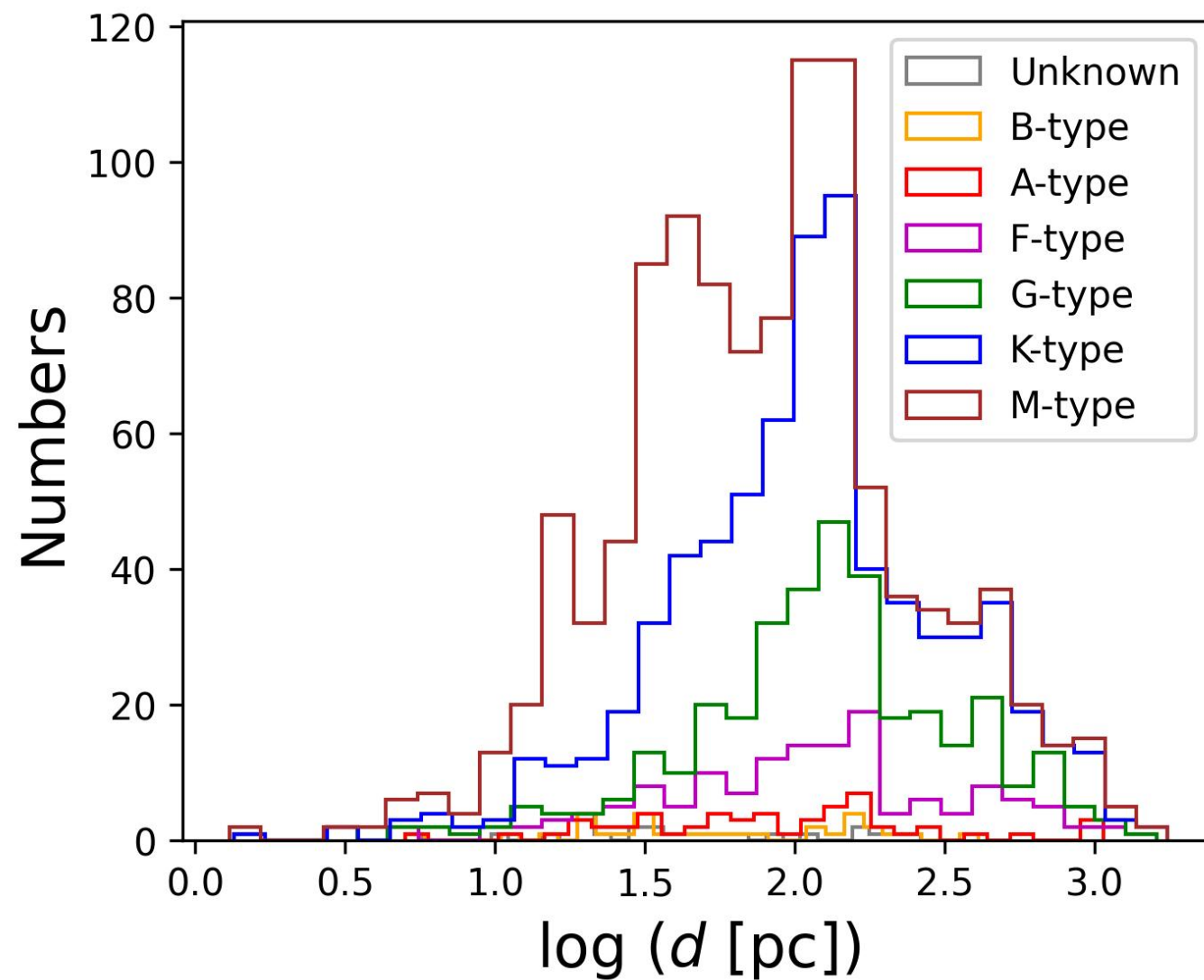


耀发恒星光谱型分布



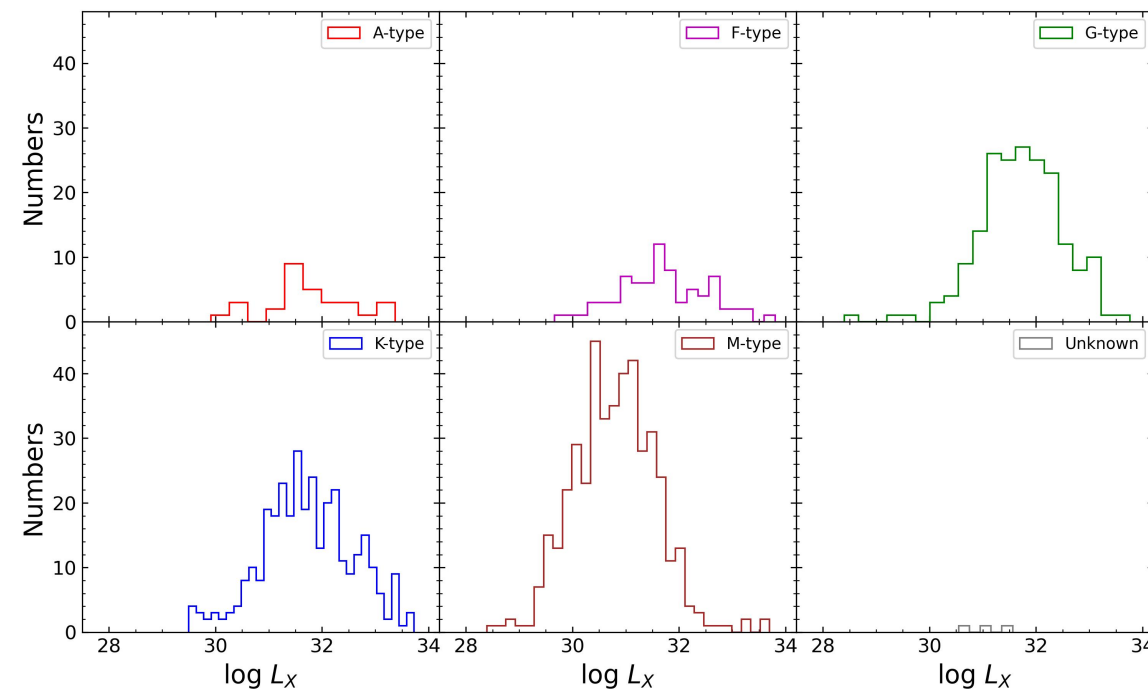
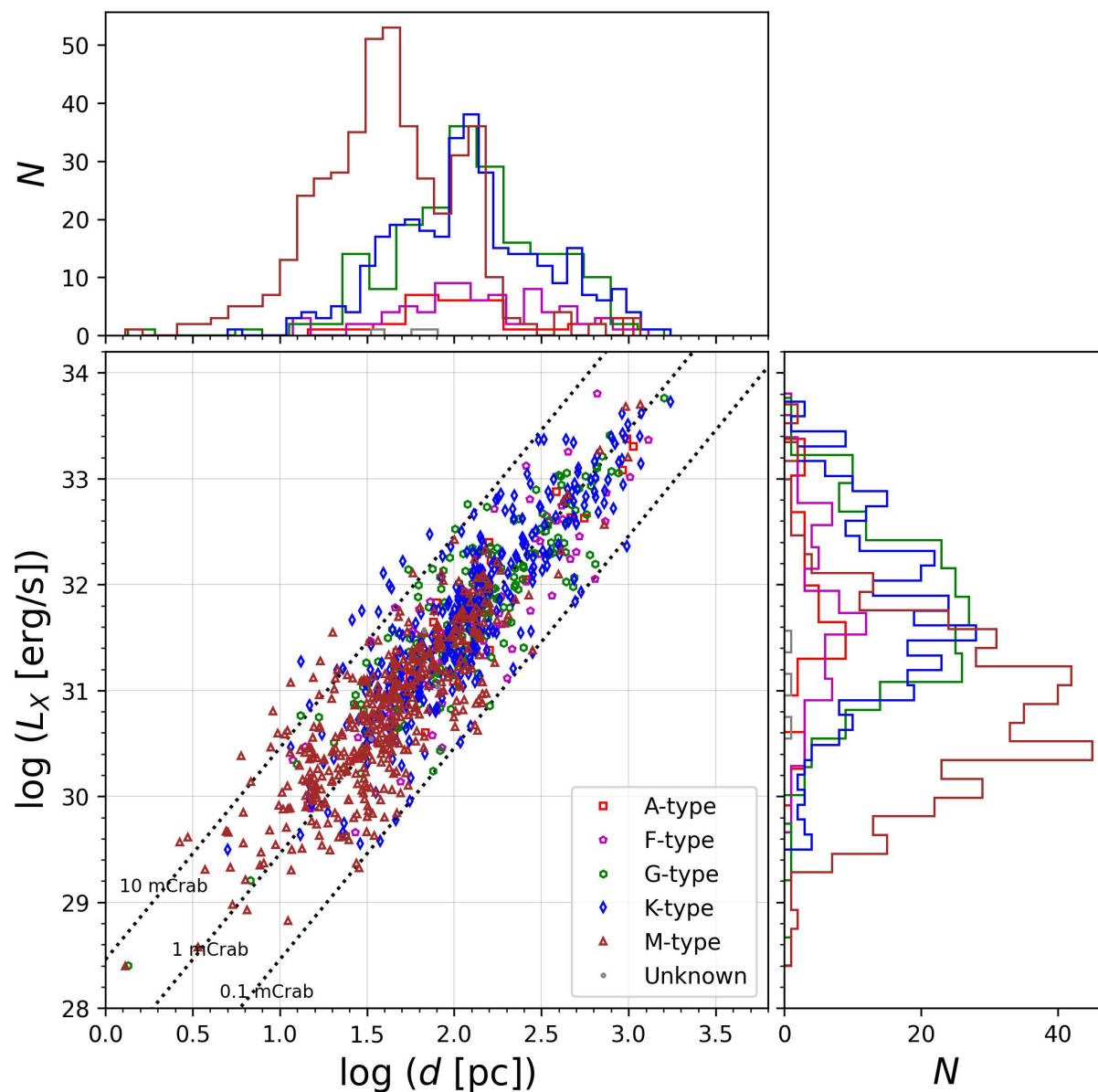
- 按测光数据BP-RP对恒星进行粗略分类
- 光谱型越晚的恒星数量越多

耀发恒星距离分布



- 双峰结构
- M型星与其他光谱型恒星分布存在系统差异

耀发光度分布



光度范围: $2e28 - 6e33 \text{ erg/s}$

$\geq 1e32 \text{ erg/s}$: 252 stars

$\geq 1e33 \text{ erg/s}$: 46 stars

1 mCrab = $2.4E-11 \text{ erg/cm}^2/\text{s}$

有后随观测的恒星耀发



No.	Star	FXT	Swift	NICER	Optical Obs
1	Gaia DR3 5406245351777015040	○			
2	TYC 6079-93-1	○			
3	2MASS J12184187-0609123				○
4	HD 315687	○			
5	UPM J2224-5826	○			
6	Gaia DR3 58995295616014848	○			
7	RX J0106.4-6346	○			
8	ASAS J164602+0718.9		○		
9	LP 924-17		○		
10	LP 820-19	○	○	○	○
11	ASAS J211056+3323.2	○			
12	BPM 46566	○			
13	UCAC4 334-220322	○			
14	TYC 1625-1567-1	○			
15	Gaia DR3 2445442335531658752	○			
16	RX J0123.4+1638	○			
17	TYC 614-313-1	○			
18	UCAC4 216-002052	○			
19	MCC 401		○		
20	CRTS J225211.7+095626	○			
21	Z Ind	○			
22	2MASS J18351416+2435115	○			○
23	PM J23221-0301	○			○
24	Gaia DR3 2636220758434984448	○			

No.	Star	FXT	Swift	Optical Obs
25	2MASS J23301129-0237227	○		○
26	LSPM J0251+4434		○	
27	1RXS J225817.2-110434	○		○
28	LP 402-64	○		
29	Proxima Centauri	○		
30	Gaia DR3 2534635509050352256	○		○
31	BD-19 660		○	
32	V644 Tau	○		
33	HG 7-33	○		
34	BPS CS 31060-0015	○		
35	2MASS J04443916+2224417	○		
36	V410 Tau	○		
37	PM J02146+3121		○	
38	2MASS J04492661+2730388	○		
39	RX J0413.4-0139	○		
40	ATO J079.4052-07.5576	○		
41	UCAC4 797-019583	○		
42	Cl* Melotte 20 DH 189	○		
43	LSPM J0447+1437	○		
44	TYC 5360-423-1	○		
45	V830 Tau	○		
46	ATO J062.5554+20.7139	○		
47	ATO J147.4546+48.1119	○		
48	1RXS J060224.9-163451	○		

有后随观测的恒星耀发



No.	Star	FXT	Swift
49	ATO J042.4338+49.2939		○
50	Gaia DR3 176718460915520896	○	
51	RX J0429.3-3124	○	
52	UCAC4 545-027698	○	
53	RX J0534.6-0234	○	
54	LAMOST J05346.93+310945.3	○	
55	LP 844-28/EP J0855.3-2351	○	
56	PM J16196+7655	○	
57	RX J0448.7+1003	○	
58	RX J0916.1+0153	○	
59	UCAC4 185-006985	○	
60	UPM J1217-3644	○	
61	G 118-68	○	
62	EP250204a	○	
63	2MASS J09180165-5452332	○	
64	2MASS J23473680+2702068		○
65	2MASS J01292678+8059457	○	○
66	ATO J076.8002+24.6211		○
67	RX J0523.0-0850	○	
68	Gaia DR3 786082236117965440	○	
69	UCAC4 188-052100	○	
70	V1217 Cen	○	
71	EP J0422.0+2826	○	
72	Gaia DR3 5563280076339409920	○	

No.	Star	FXT	Swift	NICER
73	LP 378-313	○		
74	TYC 4136-261-1	○		
75	ATO J076.8002+24.6211		○	
76	TYC 247397-986-1	○		
77	Gaia DR3 6093234407369953024	○		
78	2MASS J15504650+3054070	○		
79	Gaia DR3 5468643667425957632	○		
80	BD+03 2186	○		
81	LP 368-98	○		
82	ASAS J145837-3915.1	○		
83	CPD-56 965	○		
84	12 Mus	○		
85	TYC 5423-589-1	○		
86	Gaia DR3 4443461705748132608		○	
87	Gaia DR3 4094761312514742656	○		
88	UCAC4 234-081120	○		
89	CRTS J150851.0+032702	○		
90	CF Oct	○	○	○
91	LP 323-169	○		
92	WISEA J122501.45-521614.6	○	○	
93	Gaia DR3 6277472183485764992	○		

有FXT自主后随的耀发：共44例

2MASS J12184187-0609123

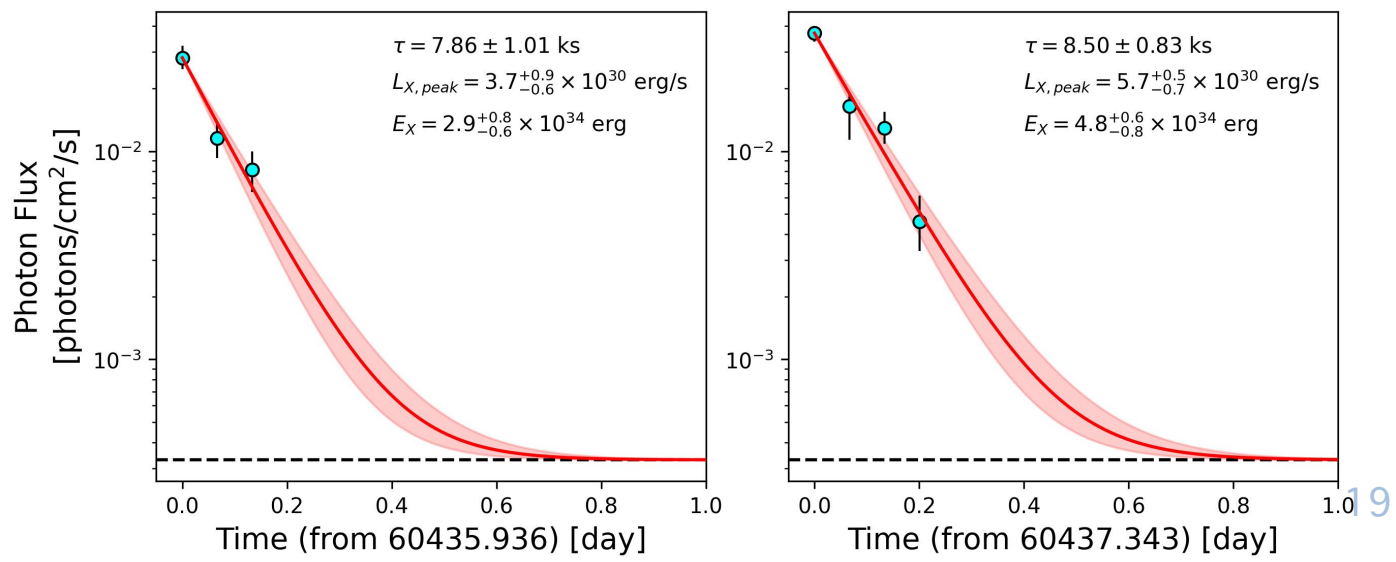
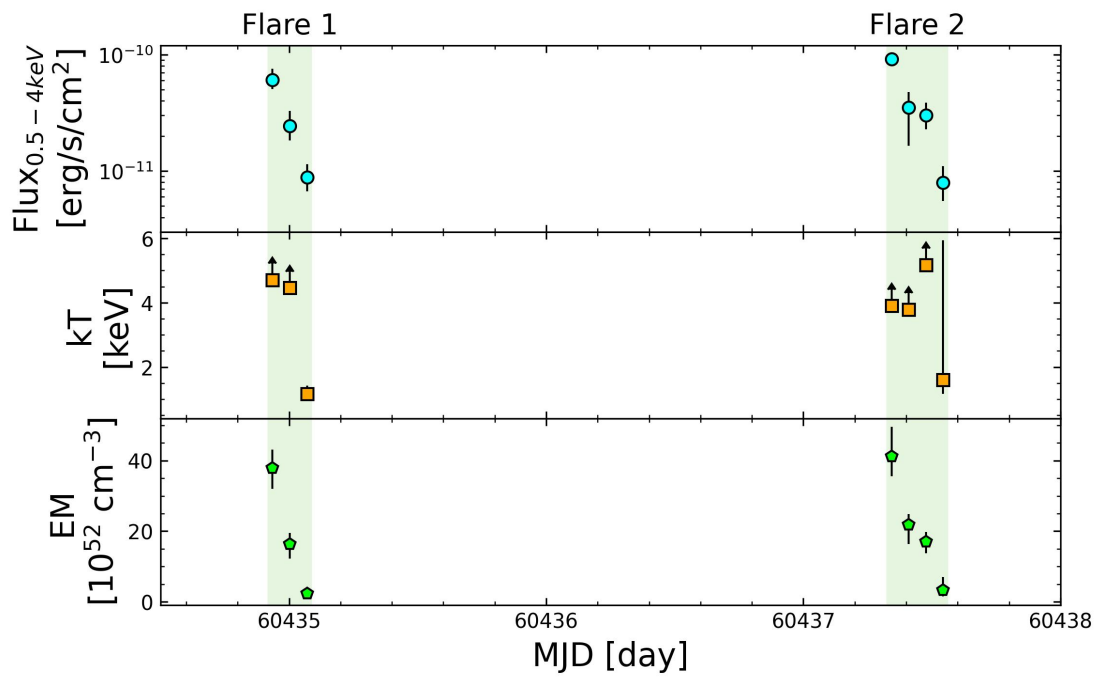


- 3天内探测到2次超级耀发

Table 2. Properties of 2MASS J12184187-0609123

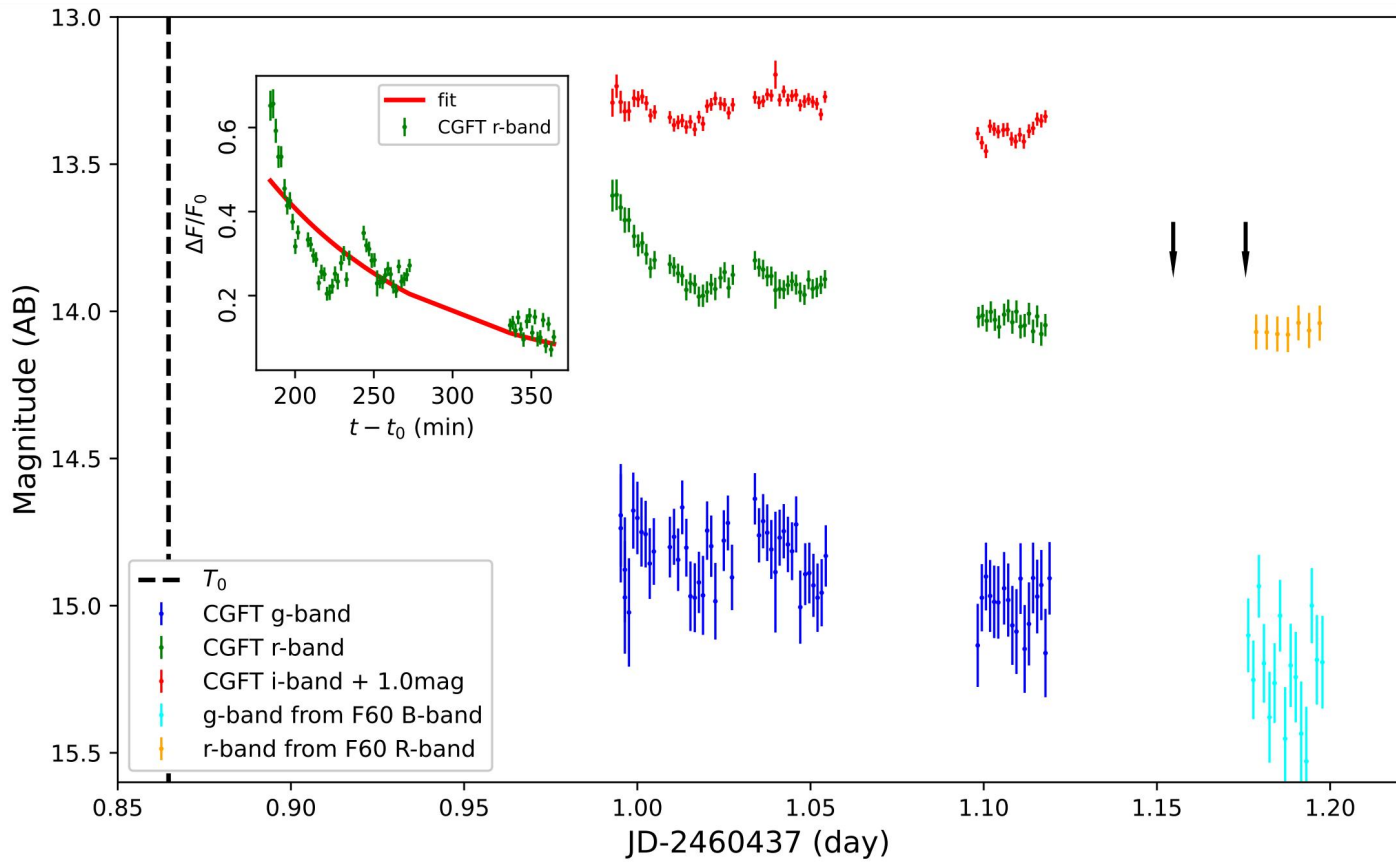
Property (1)	Value (2)
Gaia DR2 ID	3584656640992625280
G-band (mag)	13.205 ± 0.003
GBp-GRp (mag)	3.026 ± 0.007
B-band (mag)	16.34
R-band (mag)	13.43
Distance (pc)	22.64 ± 0.05
M_G (mag)	11.43
S.p. Type	M4
T_{eff} (K)	3168.0 ± 157.0
$R_\star$ ($R_\odot$)	0.256 ± 0.008
$M_\star$ ($M_\odot$)	0.227 ± 0.020
$\log(g/\text{cm s}^{-2})$	4.98 ± 0.01

Wang+ 2024





- 光学快速后随，发现H α 线移动



H α 发射线蓝移的2种解释:

1. 色球蒸发 $\rightarrow M_{PL} \sim 1.2 \times 10^{18} \text{ g}$
2. 日珥爆发 $\rightarrow 7 \times 10^{15} \text{ g} < M_p < 7 \times 10^{18} \text{ g}$

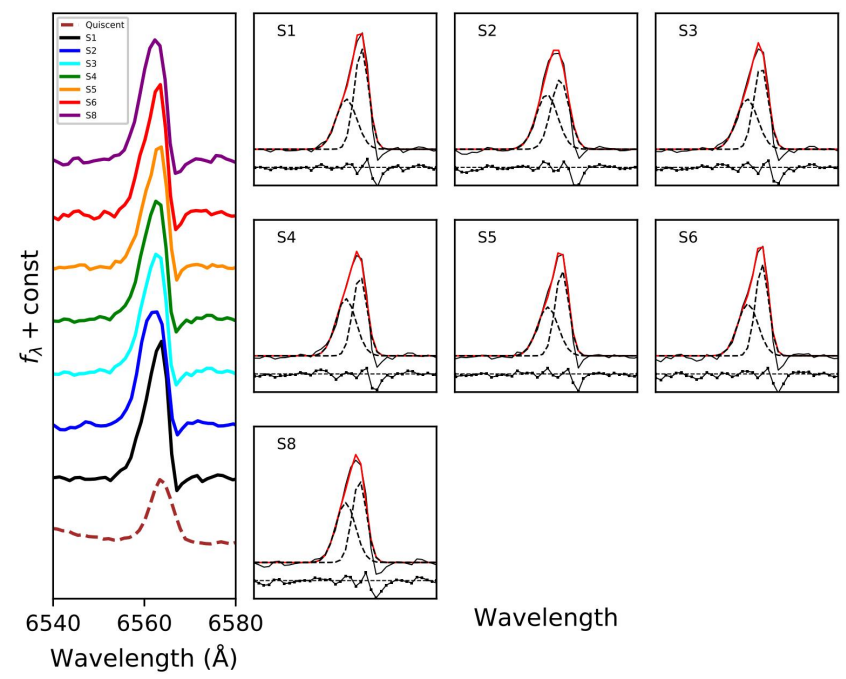


Table 3. Results of Spectral Measurements and Analysis.

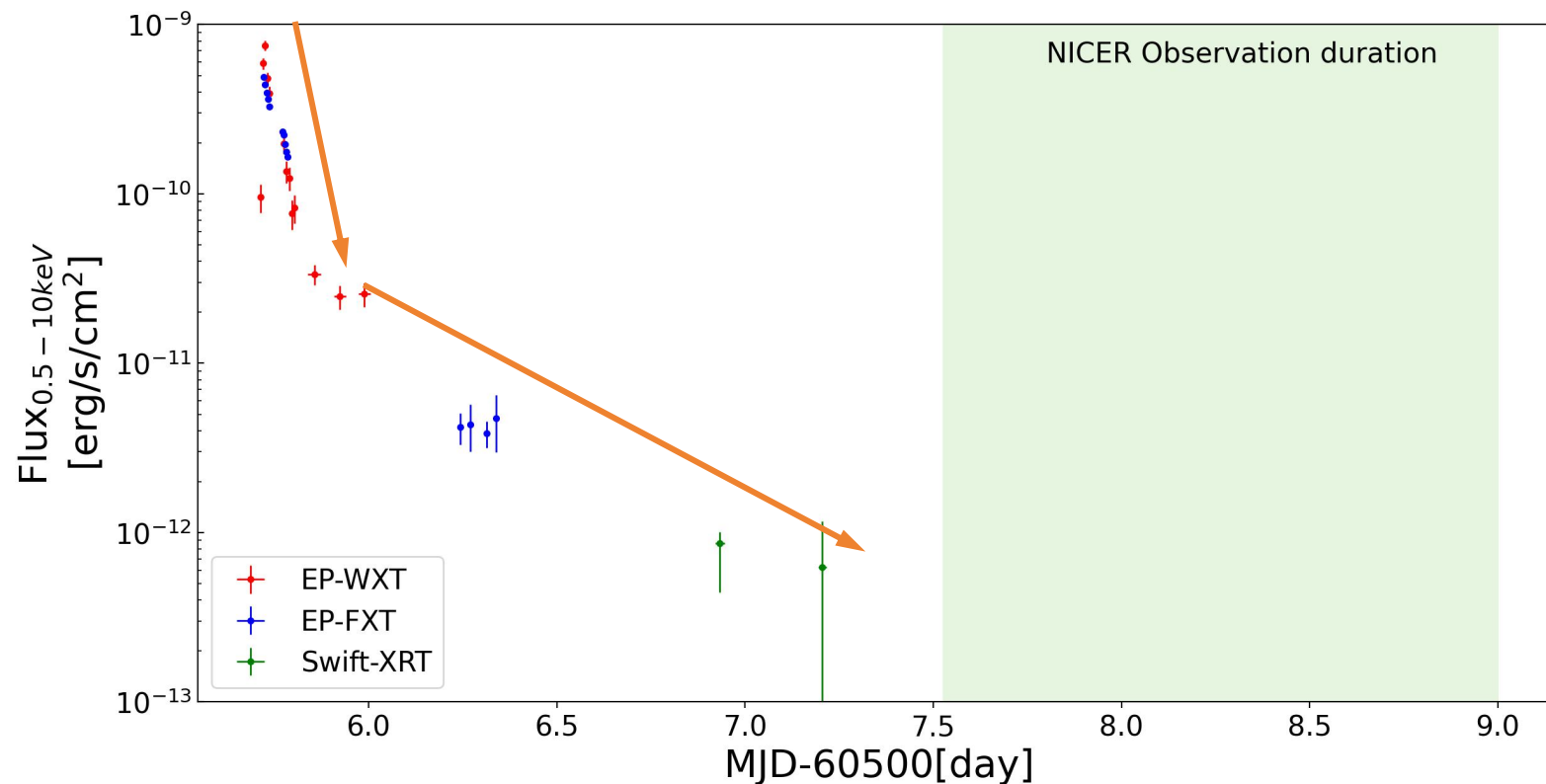
ID	$f(\text{H}\alpha_n)$ ($10^{-15} \text{ erg s}^{-1} \text{ cm}^{-2}$)	$f(\text{H}\alpha_b)$ ($10^{-15} \text{ erg s}^{-1} \text{ cm}^{-2}$)	FWHM(H α)	$\Delta v(\text{H}\alpha_b)$ (km s^{-1})	V_{max}
(1)	(2)	(3)	(4)	(5)	(6)
1	86.6 ± 1.2	63.5 ± 2.6	200	-150	-240
2	75.6 ± 8.0	71.6 ± 7.0	210	-130	-250
3	74.7 ± 3.7	65.3 ± 4.1	210	-140	-240
4	67.9 ± 2.0	69.8 ± 1.4	190	-140	-230
5	70.4 ± 2.0	61.1 ± 2.8	200	-140	-230
6	72.5 ± 1.6	64.2 ± 3.2	190	-140	-180
8	70.8 ± 2.5	72.0 ± 3.3	190	-140	-210

LP 820-19



- X射线流量下降4个量级

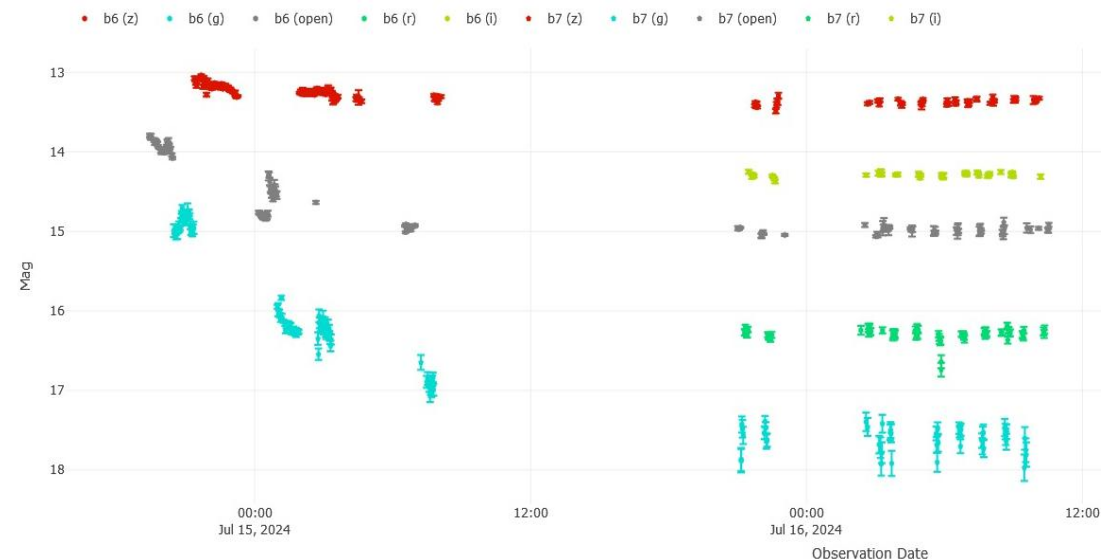
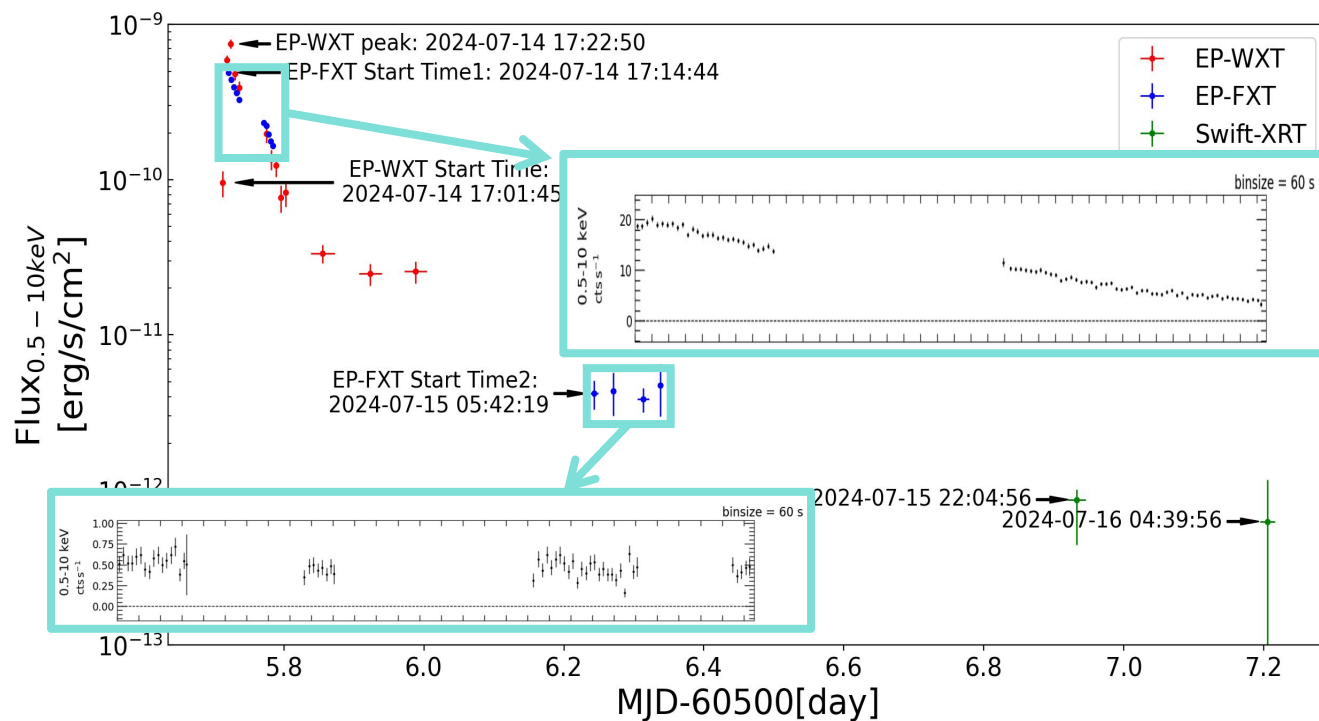
Type	High Proper Motion Star
RA, DEC	336.50 -15.30
Galactic (l, b)	44.72 -54.26
Proper motions	272.423 -53.519
Spectral type	M5.5
Distance	24.4 pc
Mag G	14.96
Teff	2799.2
log g	5.1
L_star/L_sol	1.67e-03
R_star/R_sol	0.166
M_star/M_sol	0.159



LP 820-19



- 多波段观测及时后随



BOOTES-6/DPR & BOOTES-7

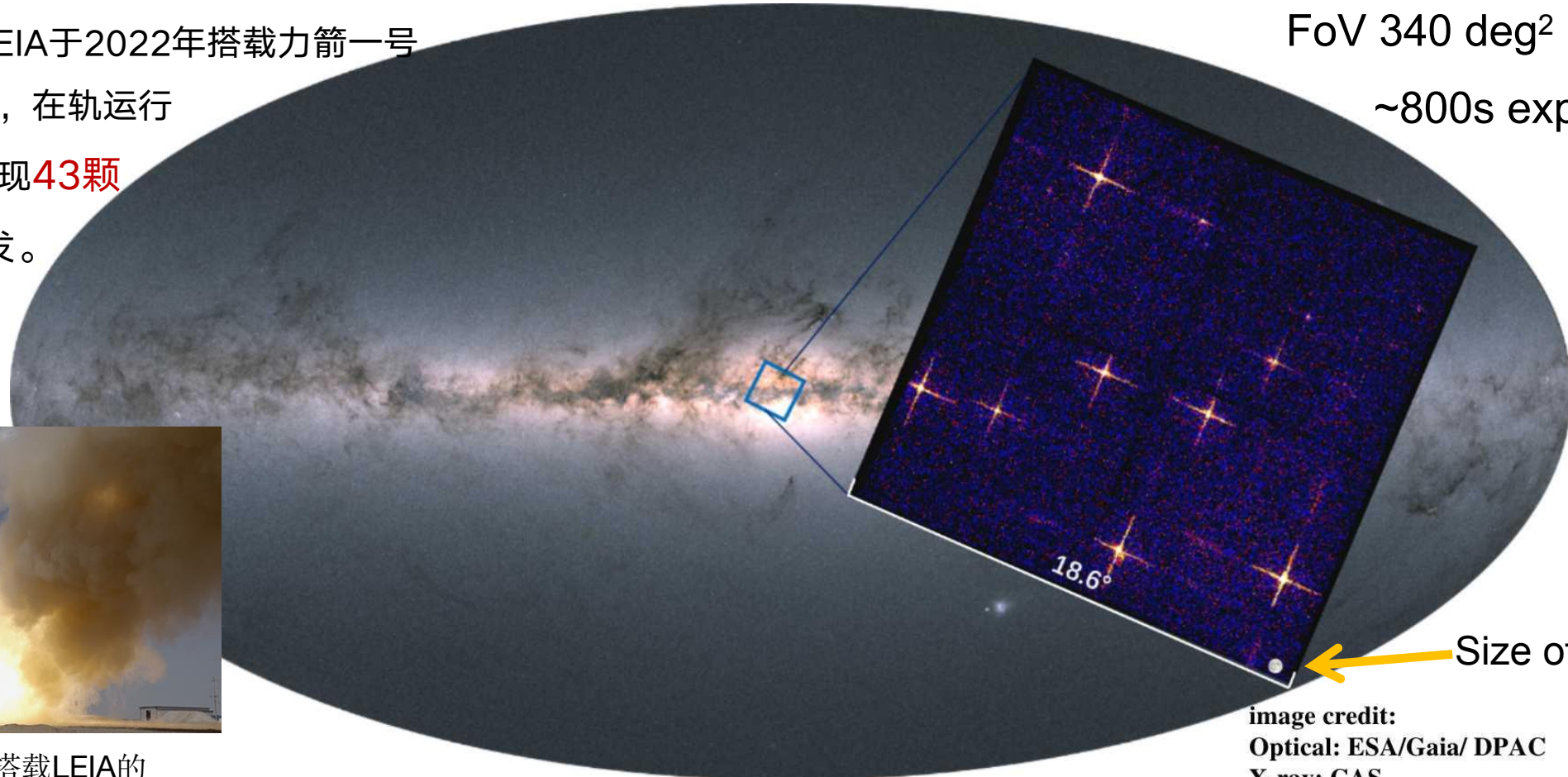
LEIA: EP探路者卫星



LEIA (Lobster Eye Imager for Astronomy) 是EP的探路者卫星，具有340平方度的大视场软X射线探测和成像能力，其科学功能可充分代表EP-WXT的12个模块之一。LEIA于2022年搭载力箭一号运载火箭成功发射，在轨运行两年多以来，已发现**43颗**恒星共计**98次**耀发。



2022年7月27日，搭载LEIA的SATech-01发射升空



FoV 340 deg²

~800s exposure

18.6°

Size of full Moon

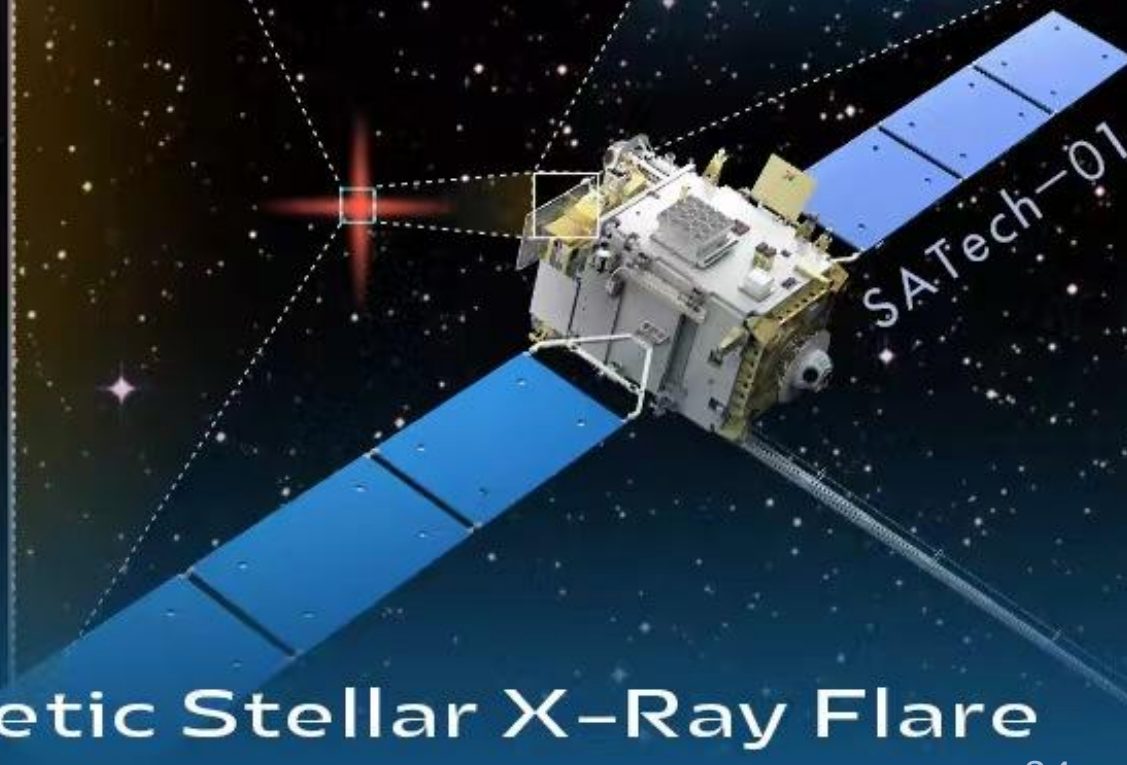
image credit:
Optical: ESA/Gaia/ DPAC
X-ray: CAS

Instrument test observations August 8-10, 2022

LXT221107A: 最强恒星耀发



LEIA



Discovery of Most Energetic Stellar X-Ray Flare

LXT221107A: 最强恒星耀发

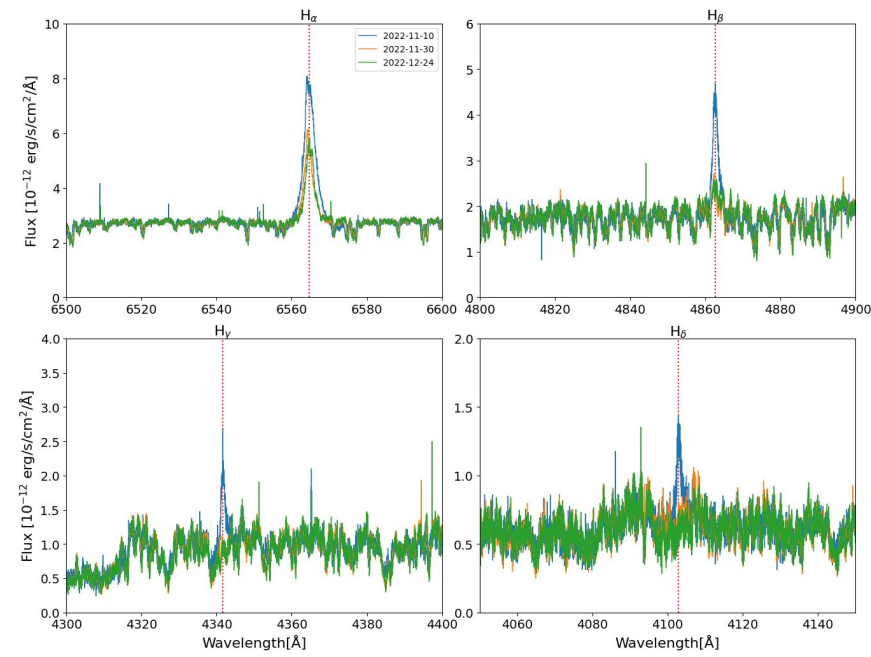
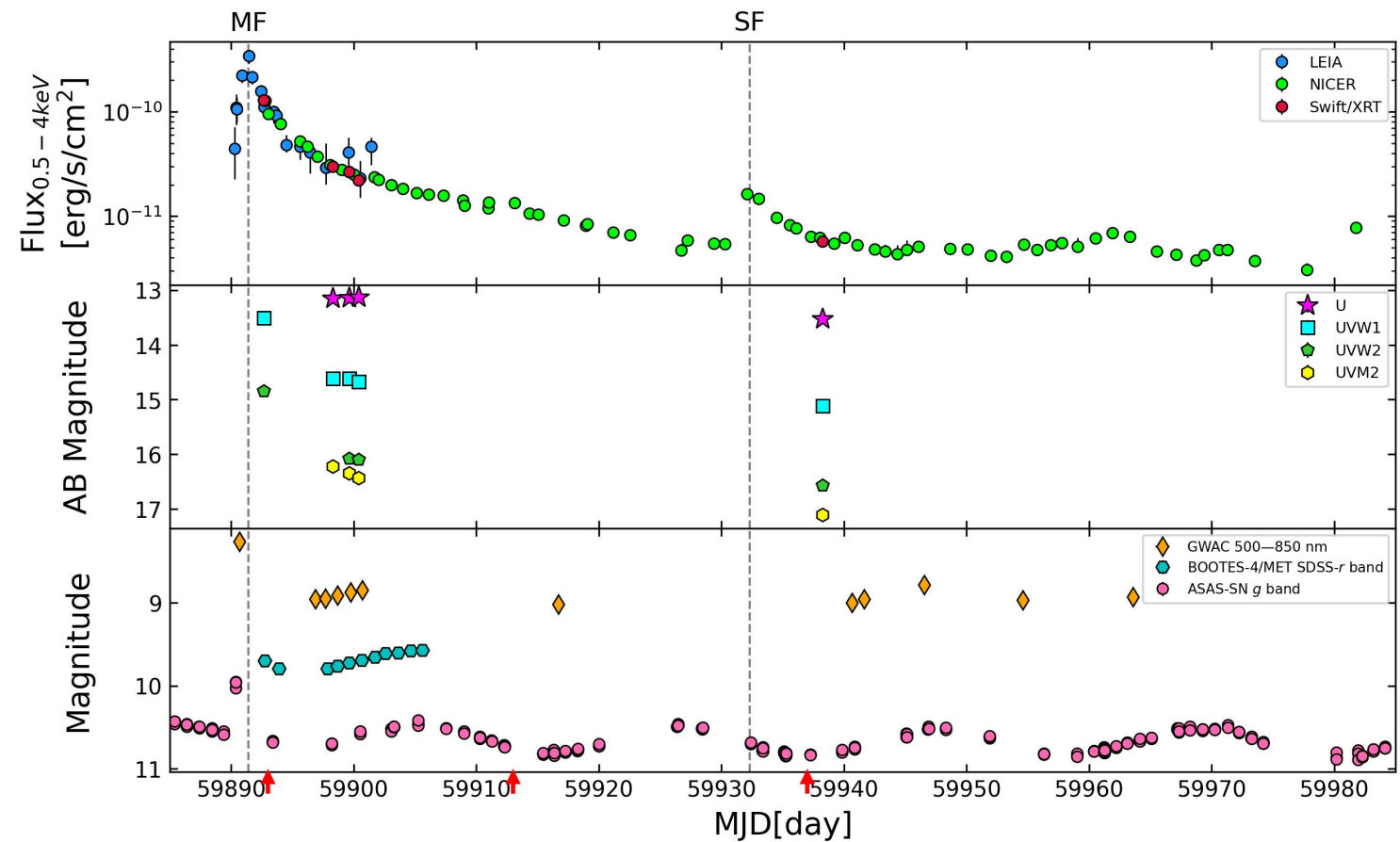


Table 1. Multi-wavelength observations on HD 251108

Instrument	Obs-Date	Waveband
LEIA	from 2022-11-07 to 2022-11-18	Soft X-ray
Swift	from 2022-11-09 to 2022-12-25	Soft X-ray/UV
NICER	from 2022-11-09 to 2023-02-06	Soft X-ray
ASAS-SN	from 2022-11-02 to 2023-02-09	optical
GWAC	from 2022-11-07 to 2023-01-19	optical
BOOTES-4/MET	from 2022-11-09 to 2022-11-22	optical
Lijiang 2.4m Telescope	2022-11-10, 2022-11-30 and 2022-12-24	optical

MF: $L_{X,p} \sim 1.1e34$ erg/s (0.5-4 keV)
SF: $L_{X,p} \sim 5.8e32$ erg/s (0.5-4 keV)

Mao+ 2025

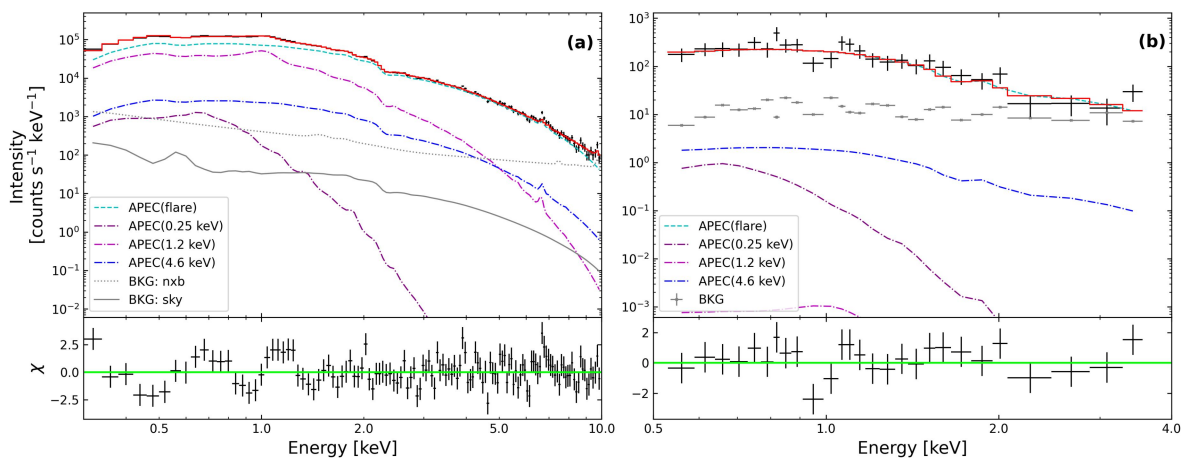
LXT221107A: 最强恒星耀发



Quiescent Spectrum (MJD>59940): 3T apec Flaring Spectra: 3+1T apec

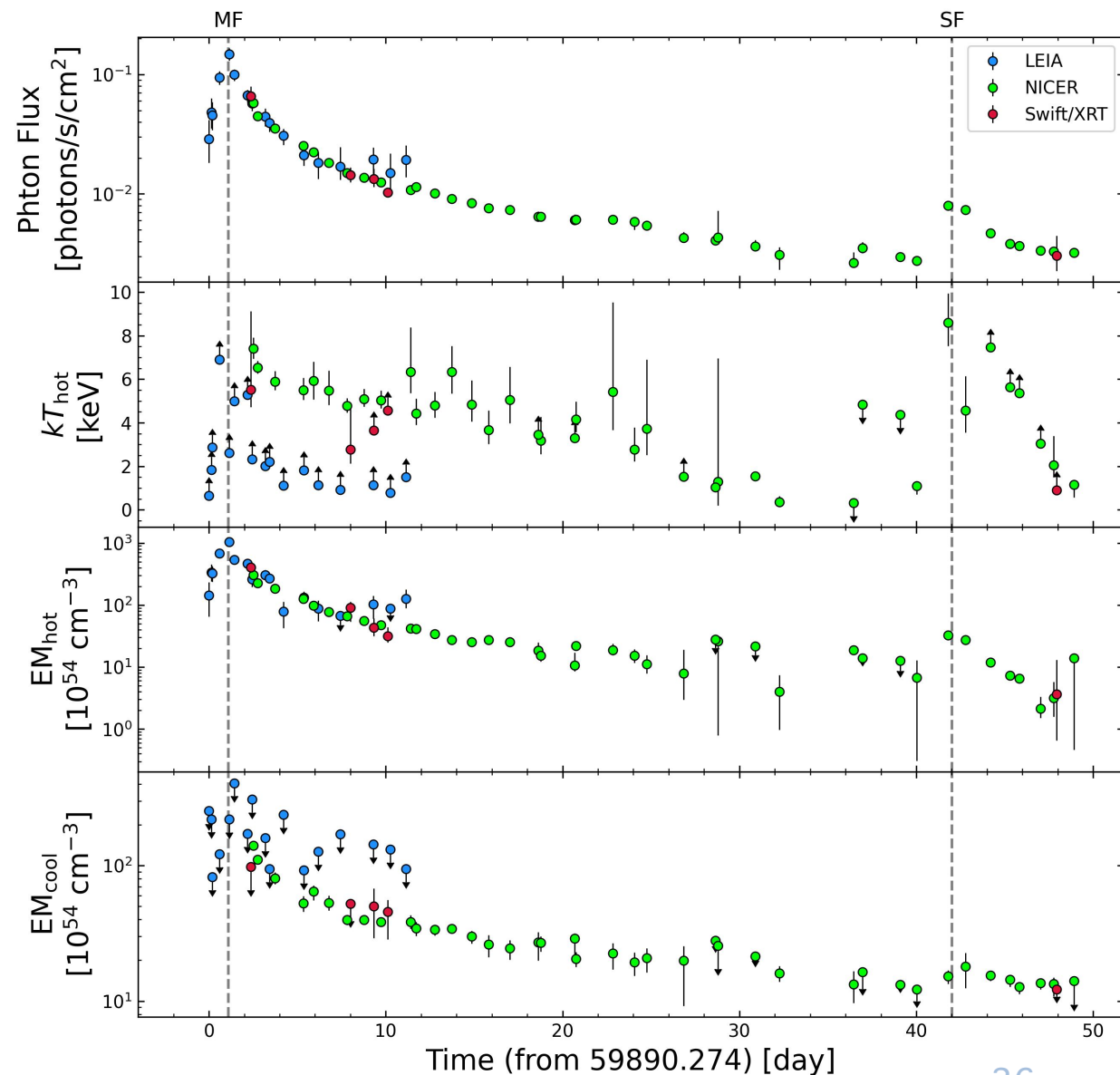
Table 2. Best-fit Spectral parameters for quiescent-state observations.

Parameters	1T apec	2T apec	3T apec
N_H^a (10^{20} cm^{-2})	6.43	$5.9^{+0.1}_{-0.3}$	$6.7^{+1.0}_{-1.6}$
$kT_{q,1}^b$ (keV)	1.57	$1.20^{+0.03}_{-0.03}$	$1.20^{+0.03}_{-0.04}$
$EM_{q,1}^c$ (10^{54} cm^{-3})	25.0	$17.9^{+1.4}_{-1.7}$	$12.3^{+4.3}_{-2.4}$
$kT_{q,2}^b$ (keV)	-	$8.1^{+3.6}_{-2.1}$	$4.6^{+4.7}_{-0.8}$
$EM_{q,2}^c$ (10^{54} cm^{-3})	-	$6.9^{+1.0}_{-0.7}$	$9.4^{+1.5}_{-2.4}$
$kT_{q,3}^b$ (keV)	-	-	$0.25^{+0.04}_{-0.03}$
$EM_{q,3}^c$ (10^{54} cm^{-3})	-	-	$4.4^{+2.3}_{-1.3}$
Z^d ($Z_\odot$)	0.04	$0.04^{+0.01}_{-0.01}$	$0.07^{+0.02}_{-0.02}$
χ_ν^2 (d.o.f.) ^e	3.18 (128)	1.36 (126)	1.06 (124)



NICER能谱

LEIA能谱

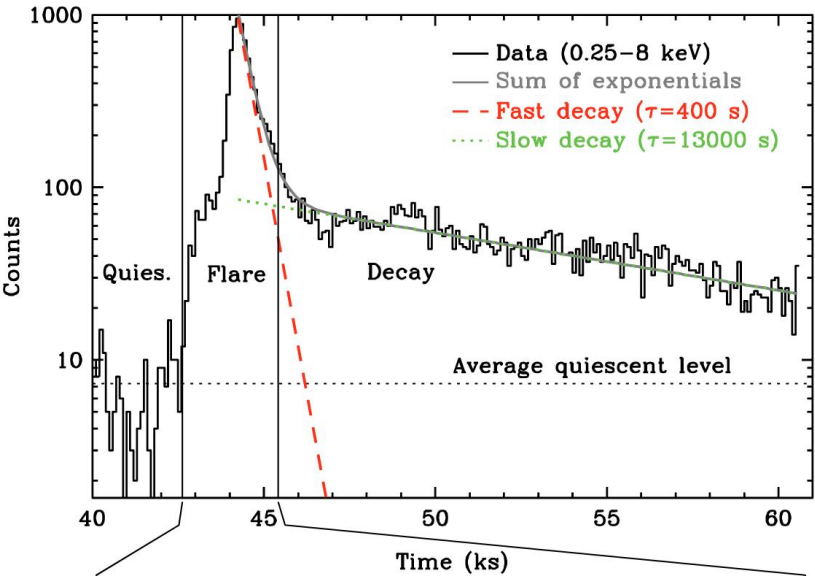


LXT221107A: 最强恒星耀发



Fast Rise and Exponential Decay (FRED) model

$$c(t) = \begin{cases} c_q & , (t < t_{ST}); \\ (c_p - c_q) \times \frac{t-t_{ST}}{t_p-t_{ST}} + c_q & , (t_{ST} \leq t < t_p); \\ (c_p - c_q) \times \exp(-\frac{t-t_p}{\tau_d}) + c_q & , (t_p \leq t). \end{cases}$$



Wargelin+ 2008

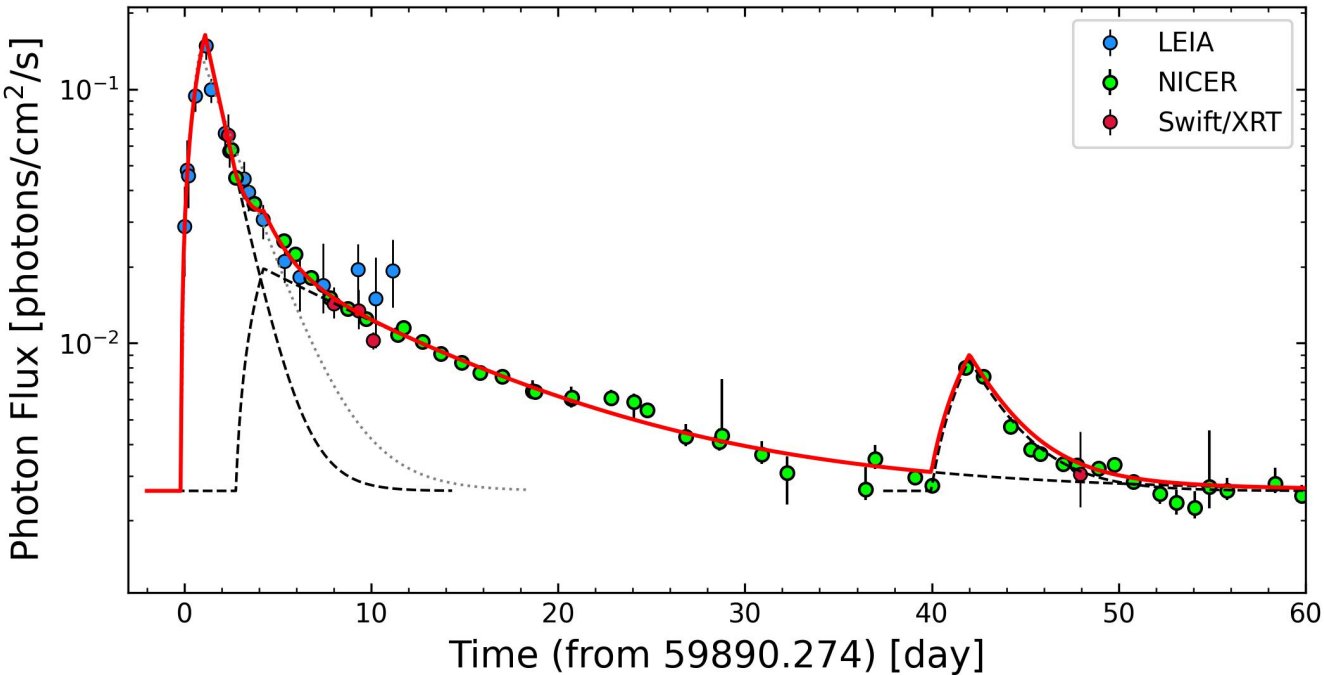


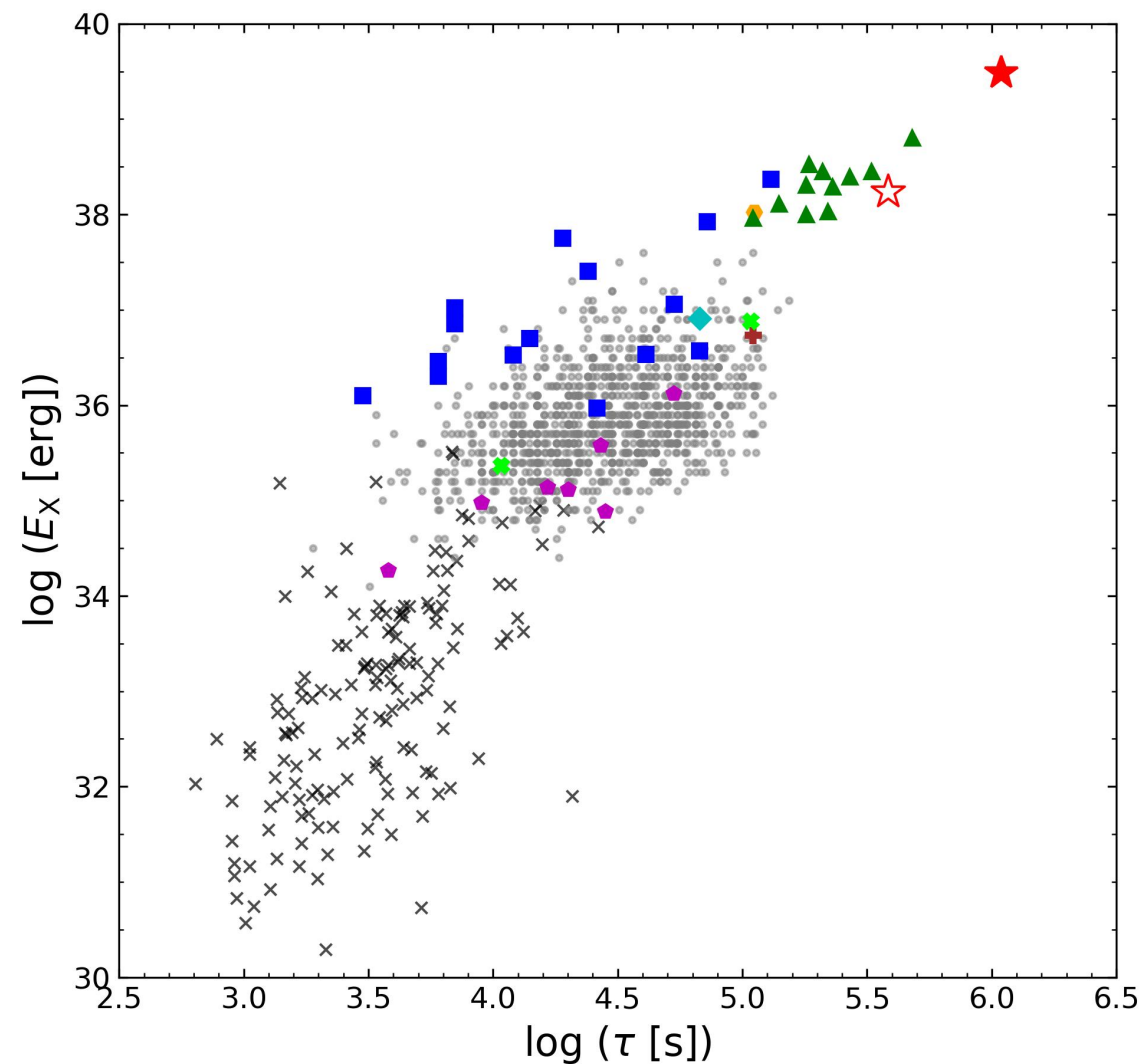
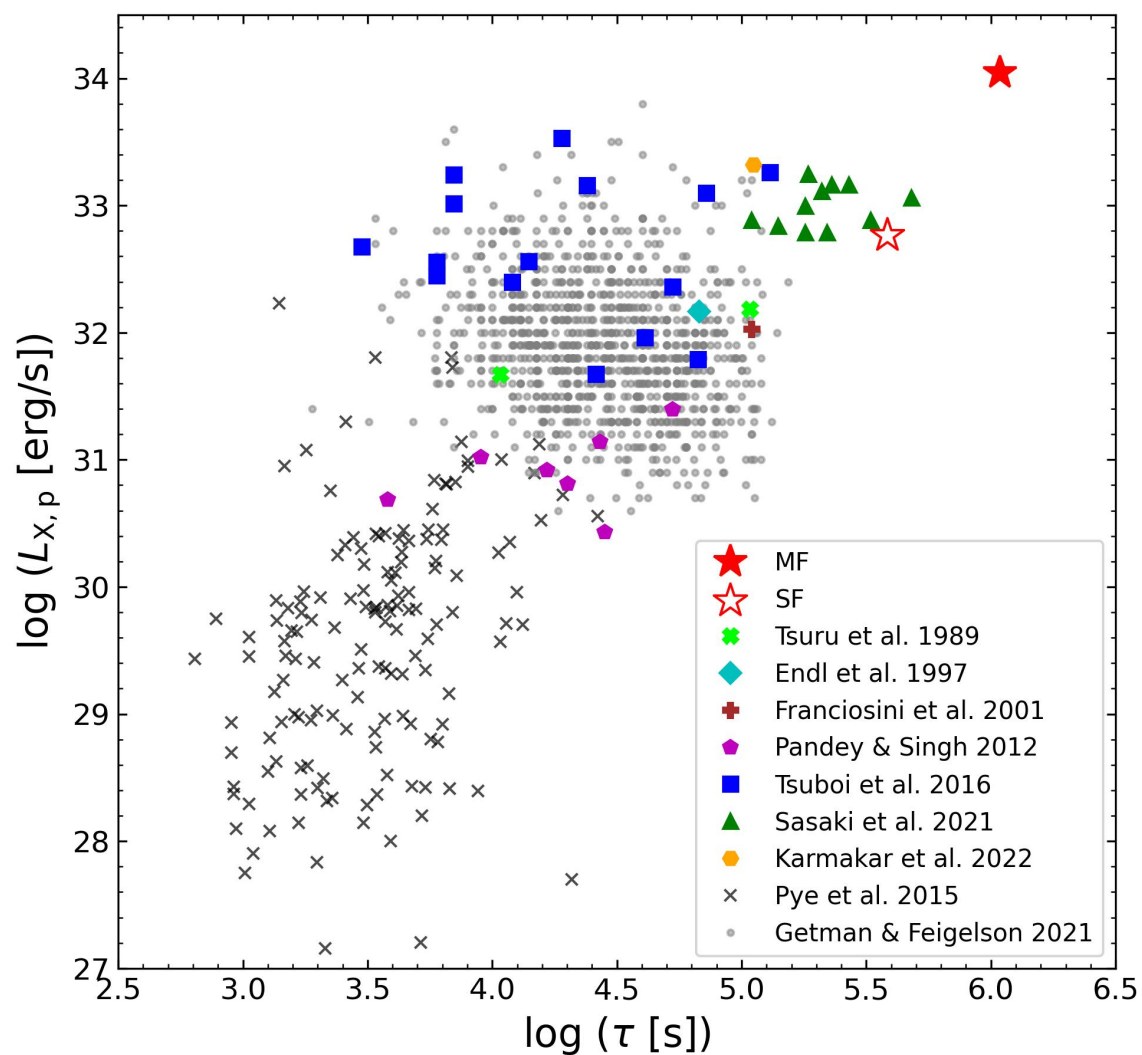
Table 3. Best-fit Parameters of the X-ray light curve

	t_p^a	c_p^b	τ_r^c	τ_d^d	$L_{X,p}^e$	f_X^f	E_X^g
	(days since MJD=59890.274)	(counts s ⁻¹ cm ⁻²)	(ks)	(ks)	(10 ³² erg s ⁻¹)	(10 ⁻⁵ erg cm ⁻²)	(10 ³⁸ erg)
MF1	1.1±0.3	0.16±0.04	112±50	109±9	110.8±29.7	6.0±1.9	18.3±5.7
MF2	4.2±0.2	0.020±0.001	128±12	869±34	13.3±1.7	4.1±1.1	12.4±1.7
SF	42.0±0.2	0.009±0.001	177±18	205±24	5.8±0.8	0.6±0.1	1.7±0.3

MF: $\tau \sim 1090$ ks, $E_X \sim 3.1e39$ erg (0.5-4 keV)

SF: $\tau \sim 382$ ks, $E_X \sim 1.7e38$ erg (0.5-4 keV)

LXT221107A: 最强恒星耀发



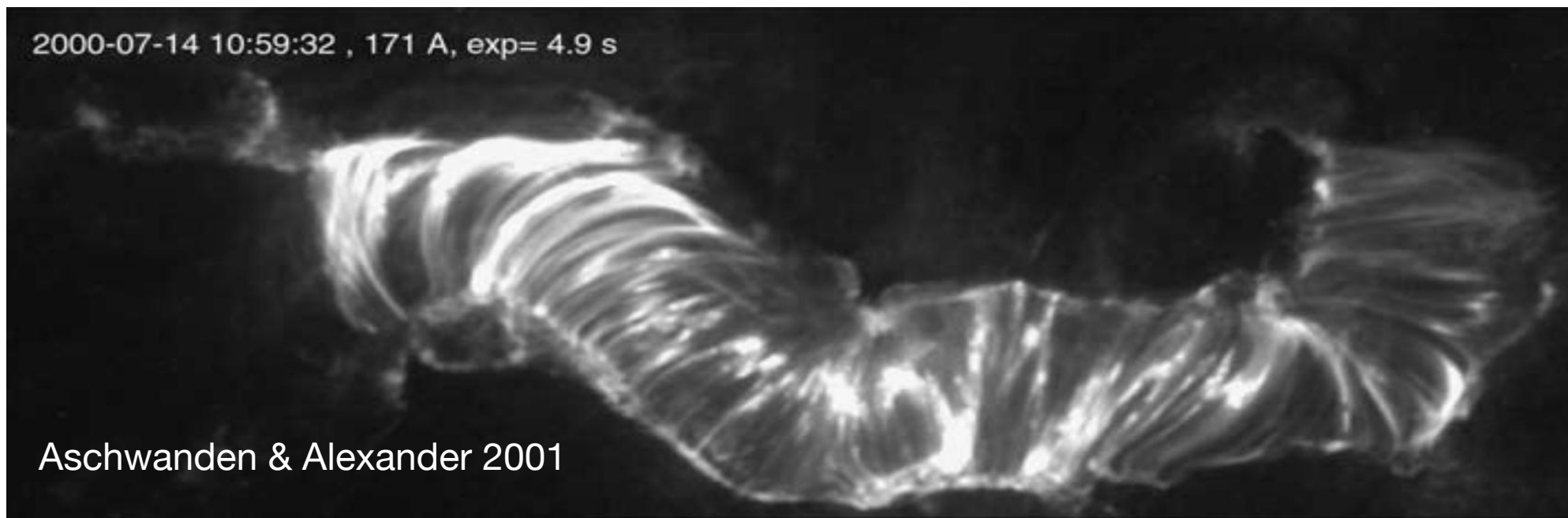
LXT221107A: 最强恒星耀发



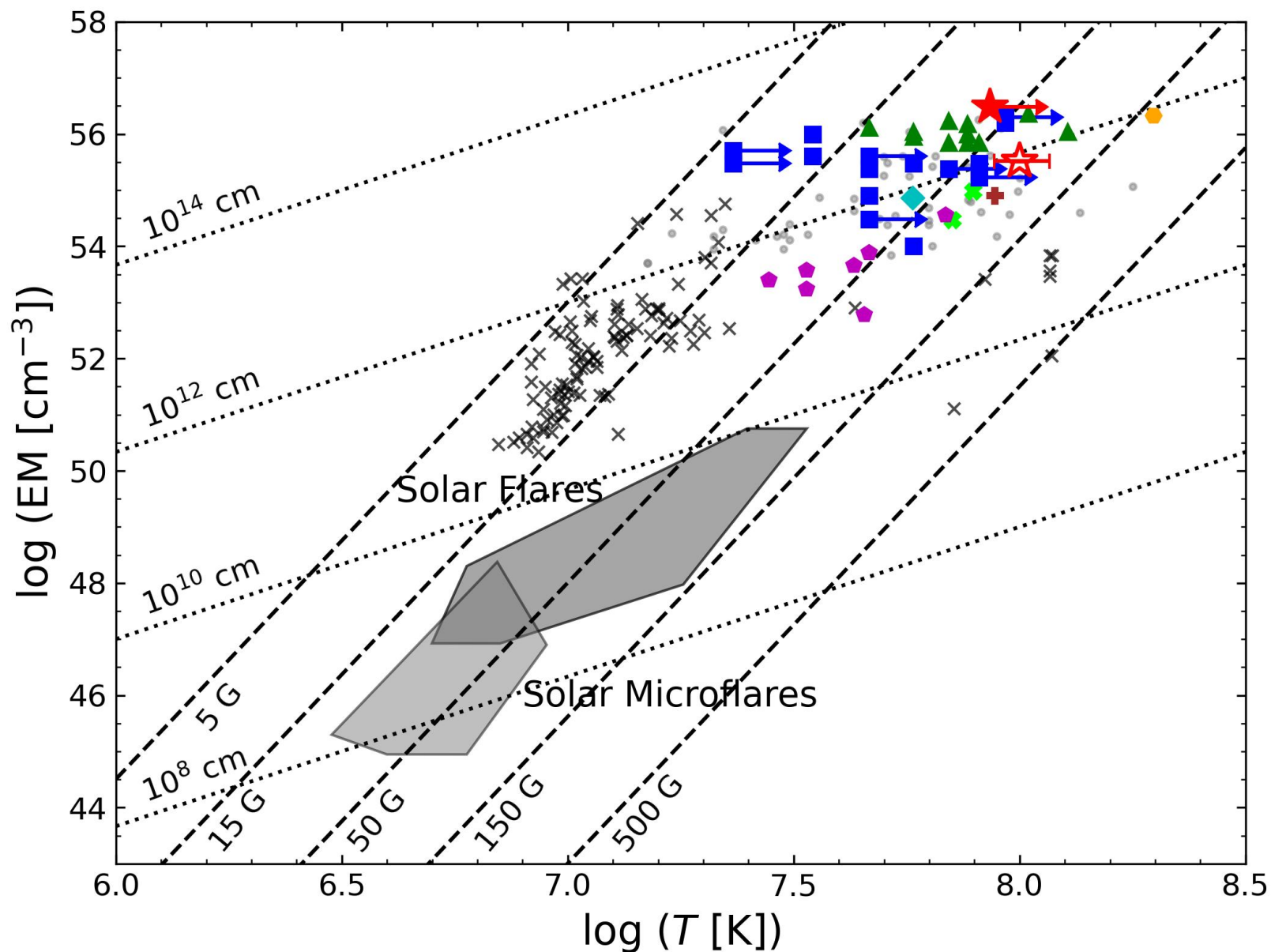
不管是太阳耀斑(Lin & Forbes 2000; Aschwanden & Alexander 2001; Lin 2002; Kashapova+ 2021), 还是恒星耀发(Reale+ 2004; Getman+ 2008; Wargelin+ 2008), 都不乏呈现出双指数下降的光变曲线。

这类耀发过程常被解释为:

- 恒星表面活跃区突然发生磁重联, 导致能量的大量释放, 最初的单个磁环被“点燃”;
- 这一最初的能量释放逐渐平息, 耀发光度迅速衰减, 表现为光变曲线中**快速下降**的部分;
- 残余的能量进而“点燃”了一系列次级磁环, 这一结构可能与太阳耀斑中观测到的磁拱廊类似 (arcade);
- 这些次级磁环的冷却相较缓慢, 进而贡献了光变曲线中**缓慢下降**的部分



LXT221107A: 最强恒星耀发



$$EM \propto B^{-5} T^{17/2}$$

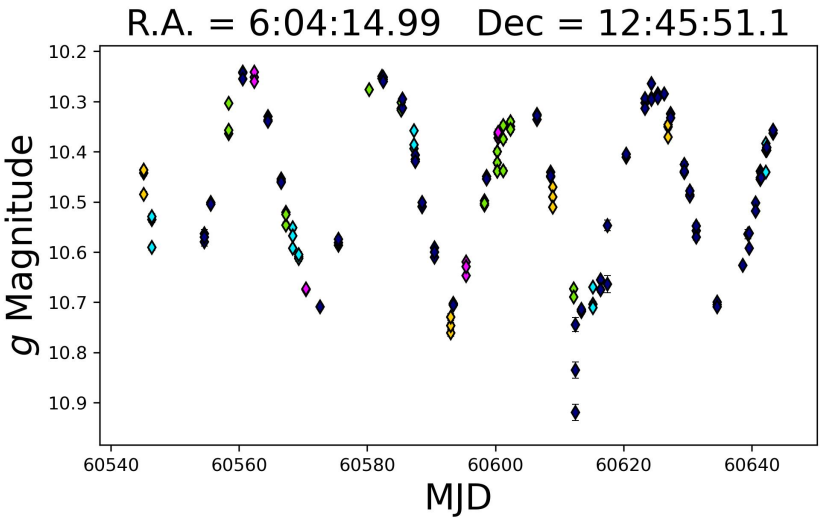
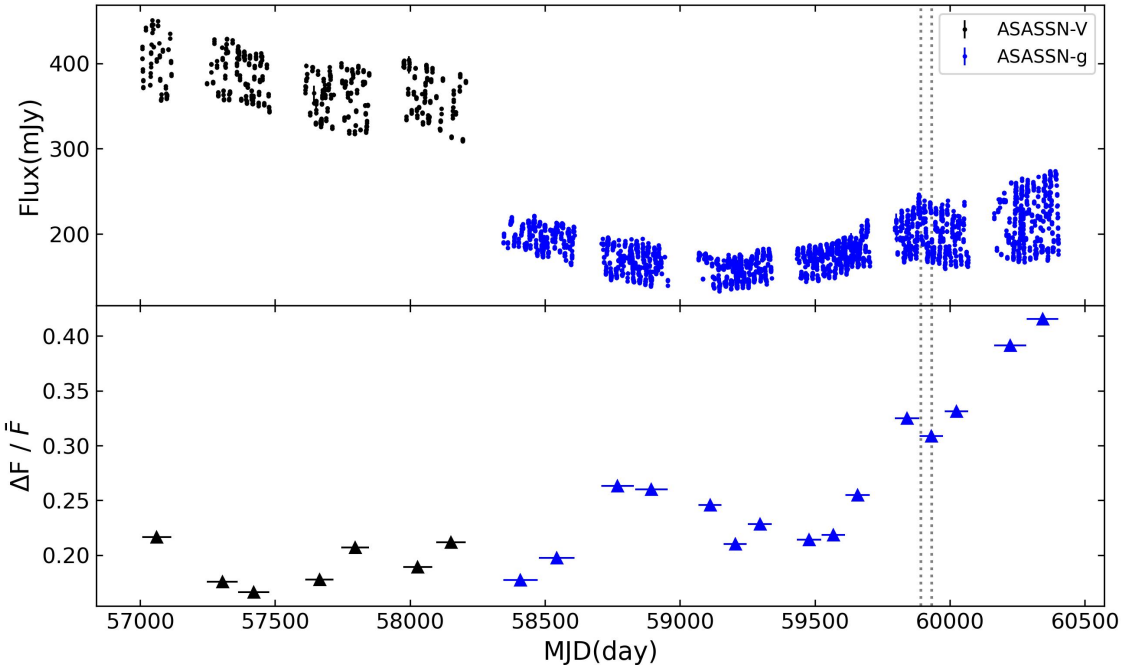
$$EM \propto L^{5/3} T^{8/3}$$

Shibata & Yokoyama 1999

LXT221107A: 恒星HD 251108

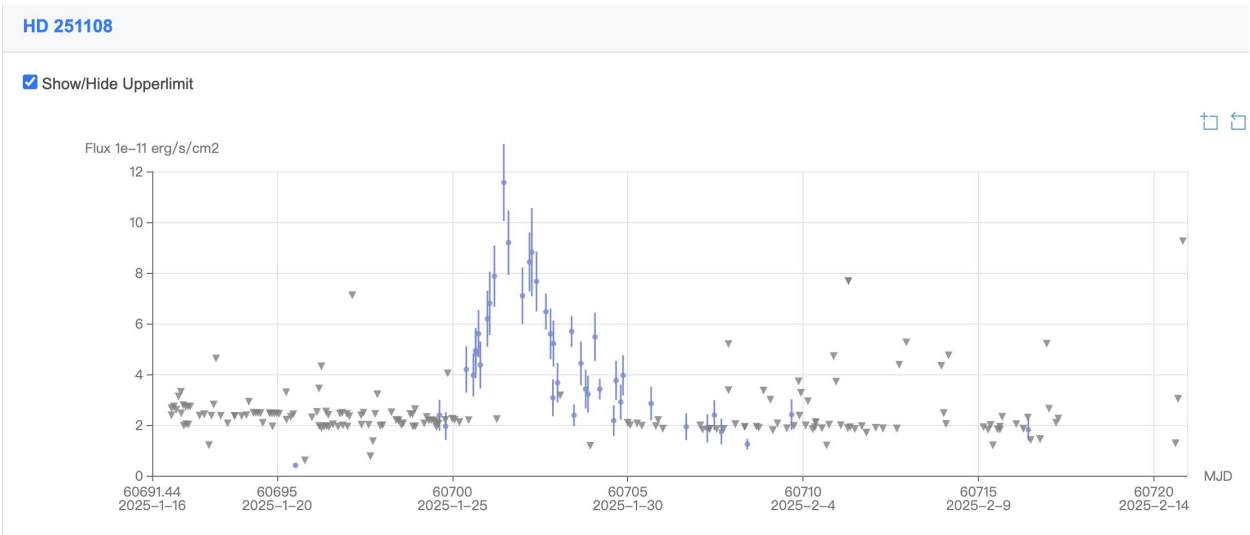


Name	HD 251108
RA, DEC	91.06, +12.76
Galactic (l, b)	196.17, -04.40
Spectral type	K2
Distance	504.7 [+4.8, -4.6] pc
Mag V	9.90 [± 0.09]
Teff	4545 [+405, -138] K
log g	2.13 [+0.17, -0.09]
M_star/M_sol	1.14 [+0.48, -0.19]



P ~21.21 d

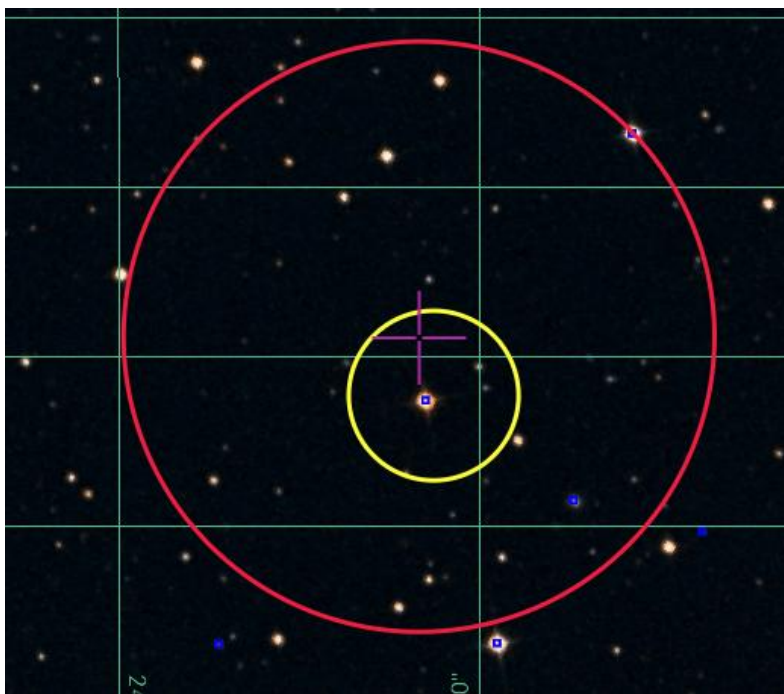
RS CVn? (Kiraga+ 2012)



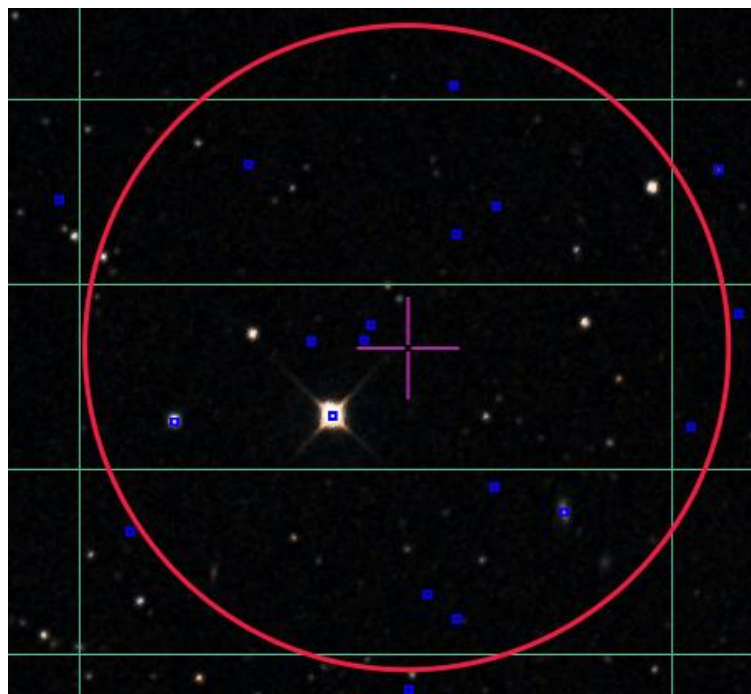
一些困难



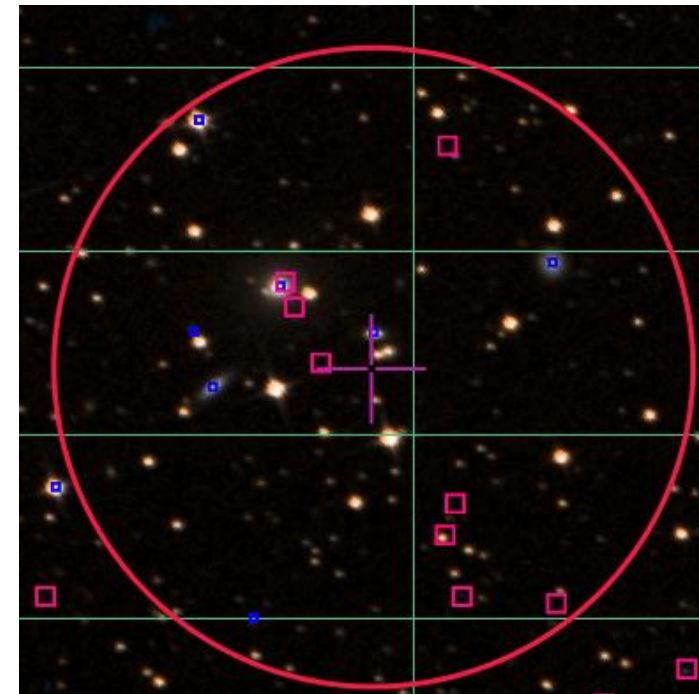
1. 许多疑似恒星耀发的探测，无法确认恒星来源



有X射线探测的恒星存在且唯一



有X射线探测的恒星不存在

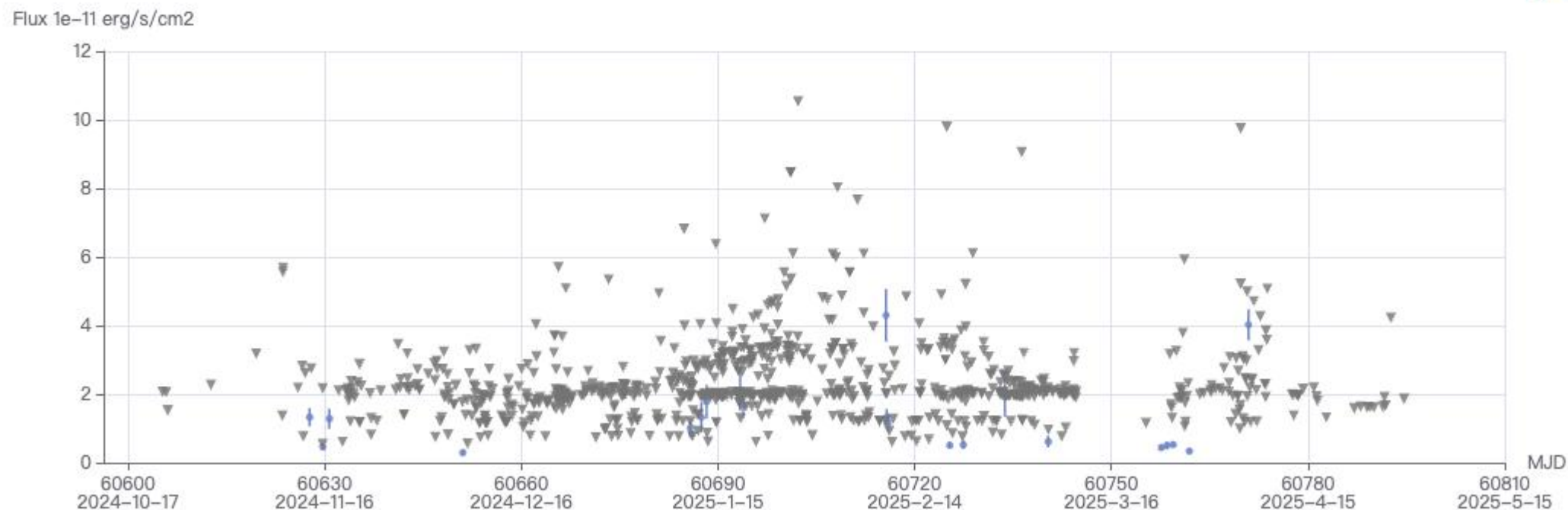


有X射线探测的恒星不唯一

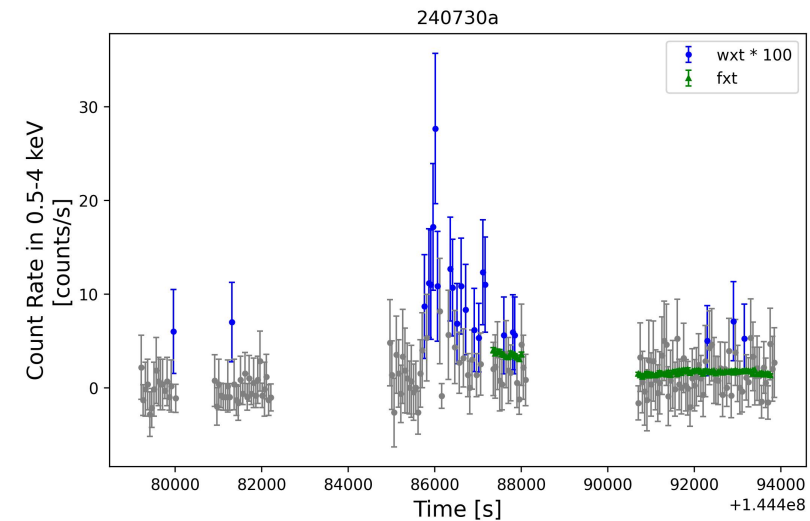
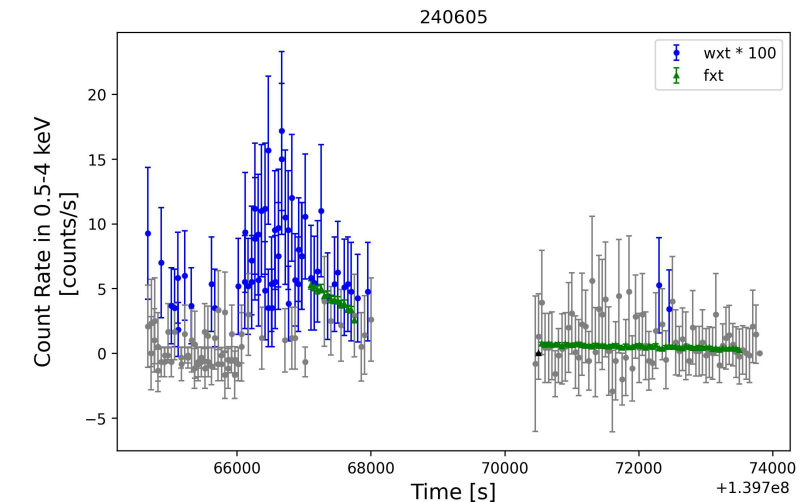
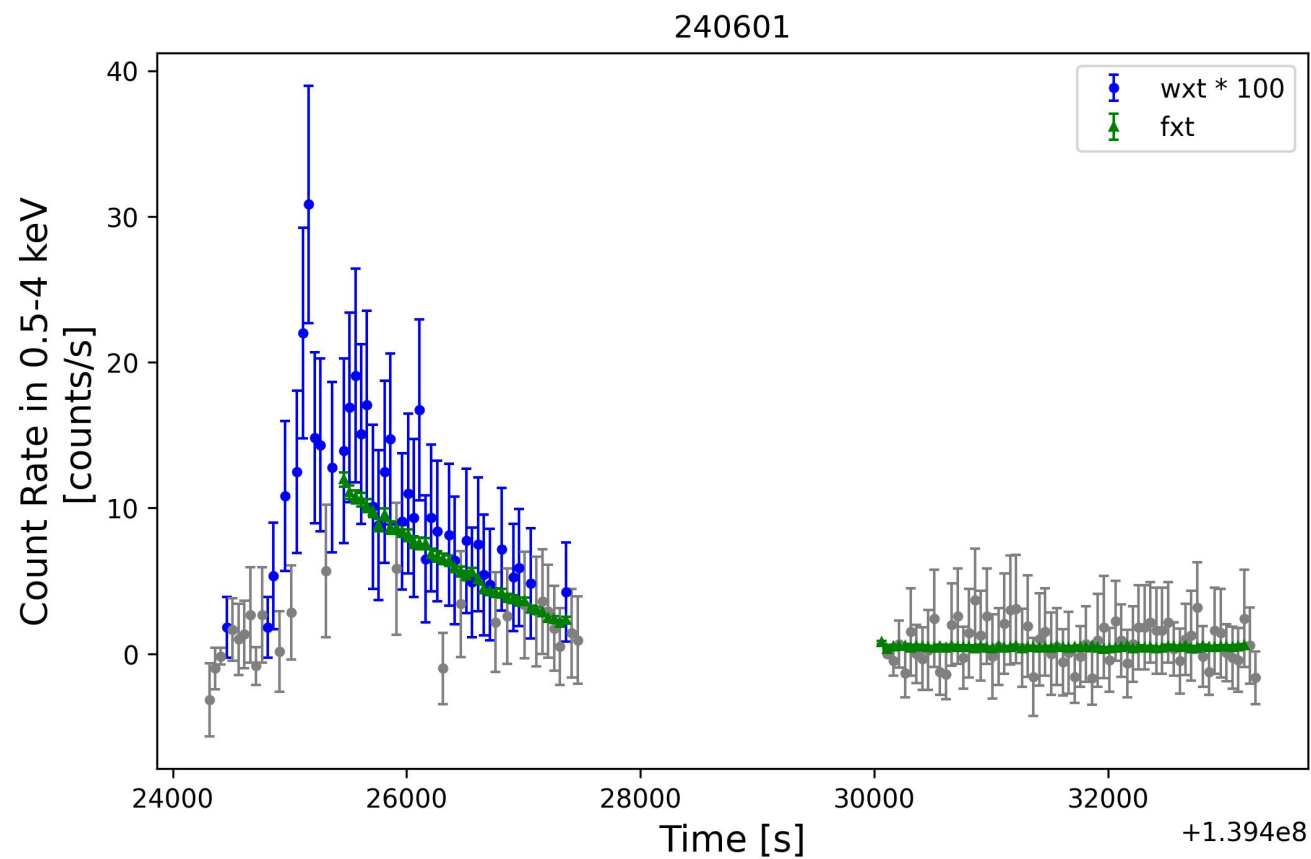
2. 大部分恒星长期持续的监测，没有X射线宁静态信息，难以对单次耀发的数据进行分析

HD 77407

☒ Show/Hide Upperlimit



3. 大部分触发的FXT后随的恒星耀发，演化较快，FXT只能探测到其下降阶段



1. 由于EP卫星具有**大视场、高灵敏度、高探测频次**的优势，对恒星耀发的探测十分有利，恒星耀发的数据正快速积累起来
2. EP升空一年多以来，共探测到来自**1101颗**恒星的耀发，主要来源于晚型星，耀发光度在 **$2e28 - 6e33 \text{ erg/s}$** 之间
3. 探测到的M型星从统计上来说距离更近，耀发光度更低
4. EP/WXT探测到的恒星耀发中共有**93例**耀发得到了其他设备的后随，其中**44例**自动触发了EP/FXT后随
5. LEIA探测到了持续40天左右的“最强”恒星X射线耀发，峰值光度达到 **$1e34 \text{ erg/s}$**
6. 目前还存在一些困难，难以识别出所有恒星耀发并进行逐个分析



LXT221107A: 根据冷却模型计算磁环参数



传导冷却主导

$$\tau_C = \frac{4 \times 10^{-10} n_e L^2}{T^{5/2}}$$

辐射冷却主导

$$\tau_R = \frac{3kT}{n_e \Lambda(T)}$$

RTV scaling law (Rosner+1978):

$$T_{\max}^2 \approx 7.6 \times 10^{-7} n_e L$$

Table B.1. Parameters of loop size and flaring plasma.

	L^a	n_e^b	V^c	β^d	p^e	B^f
	(10^{12} cm)	(10^9 cm $^{-3}$)	10^{36} cm 3		dyne cm $^{-2}$	(gauss)
MF1*	1.9 ± 0.2	5.1 ± 0.8	14.0 ± 4.6	0.56 ± 0.12	120 ± 21	54.9 ± 4.8
MF1 †	2.2	8.0	19.4	0.53	255	80.0
MF1 ‡	3.2	22.5	2.4	0.11	1442	190
MF2*	3.8 ± 0.2	1.6 ± 0.1	$61.4^{+11.5}_{-10.1}$	...	$27.3^{+3.9}_{-3.3}$	$26.2^{+1.9}_{-1.6}$
SF*	0.7 ± 0.1	$9.2^{+1.5}_{-1.4}$	0.5 ± 0.1	...	253^{+56}_{-50}	$79.8^{+8.8}_{-7.9}$

NOTE—Since the light curve of SF does not deviate from a single exponential decay, the conductive cooling phase is assumed to be skipped (Reale 2007). The cooling process of SF is then solely dominated by radiative loss.

